

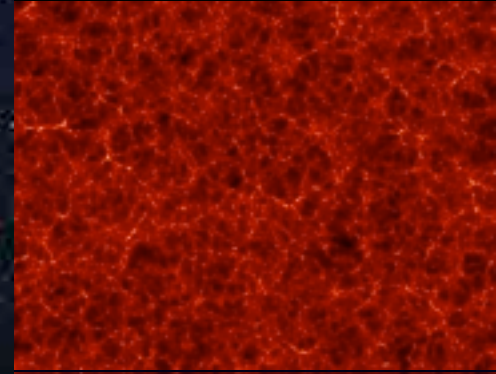
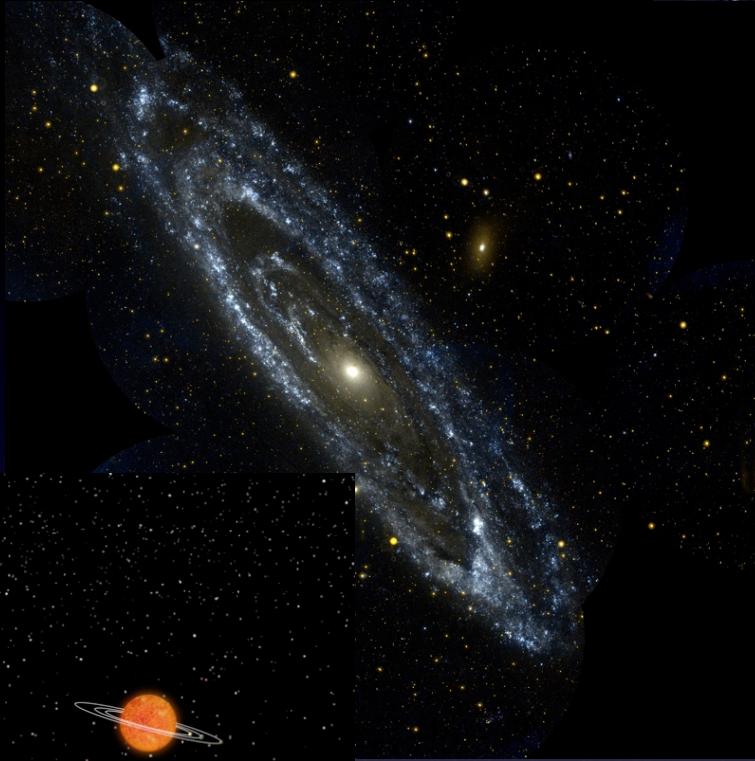
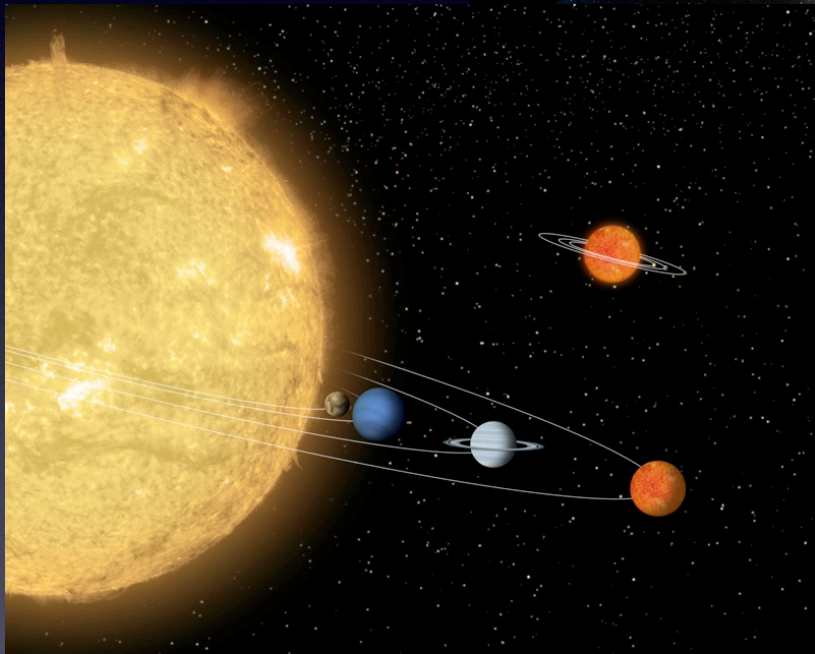
Modern Cosmology

Norman Gray
2006 November 25

Talk given at DACE Cosmologies day-school, University of Glasgow, 2006 November 25 (image credits at end)

Overview

- structure of the universe
- 'history' of western cosmologies
- cosmogeny



the universe: 3000Mpc
supercluster: 50Mpc
local group: 1Mpc=1000kpc
galaxy: 30kpc= 6×10^9 AU
solar system: 1AU= 15×10^{10} m

Scales in the universe

Cosmologies

Cosmology is a two-millennium-long retreat from anthropocentrism

(and yes, modern science does have
an *extremely* Whiggish view of history...)

Plato

somewhat hostile to
experimental science

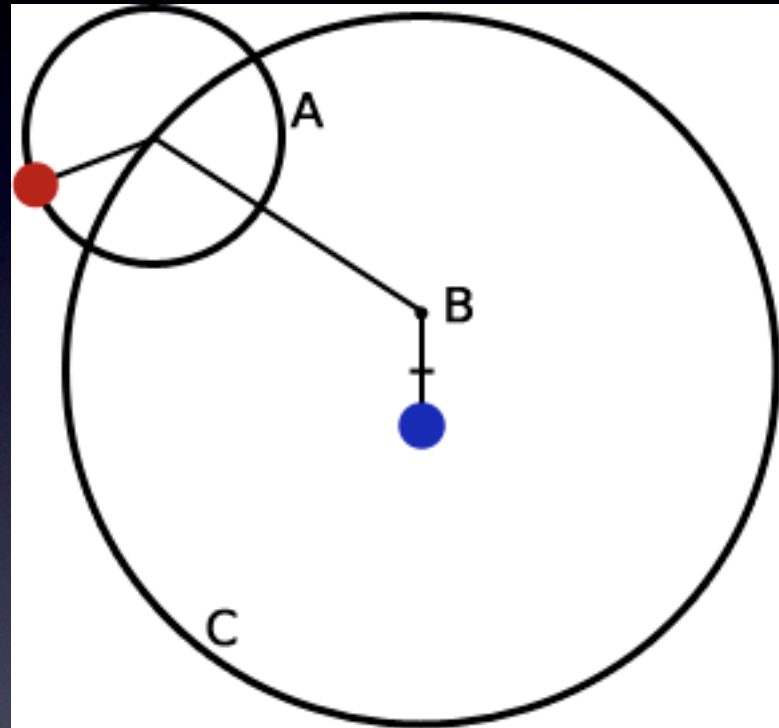
“Let us concentrate
on abstract
problems, said I, in
astronomy as in
geometry, and
dismiss the heavenly
bodies, if we intend
truly to apprehend
astronomy”

The Republic

Ptolemy

geocentric
epicycles

usable for producing almanacs
sky-geometry, divorced from sky
four elements, plus quintessence



Copernicus

heliocentric
still lots of epicycles
no *better* at ephemerides



p3c

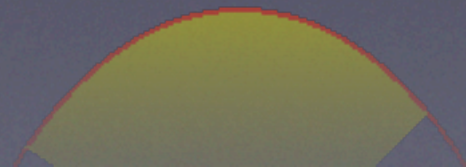
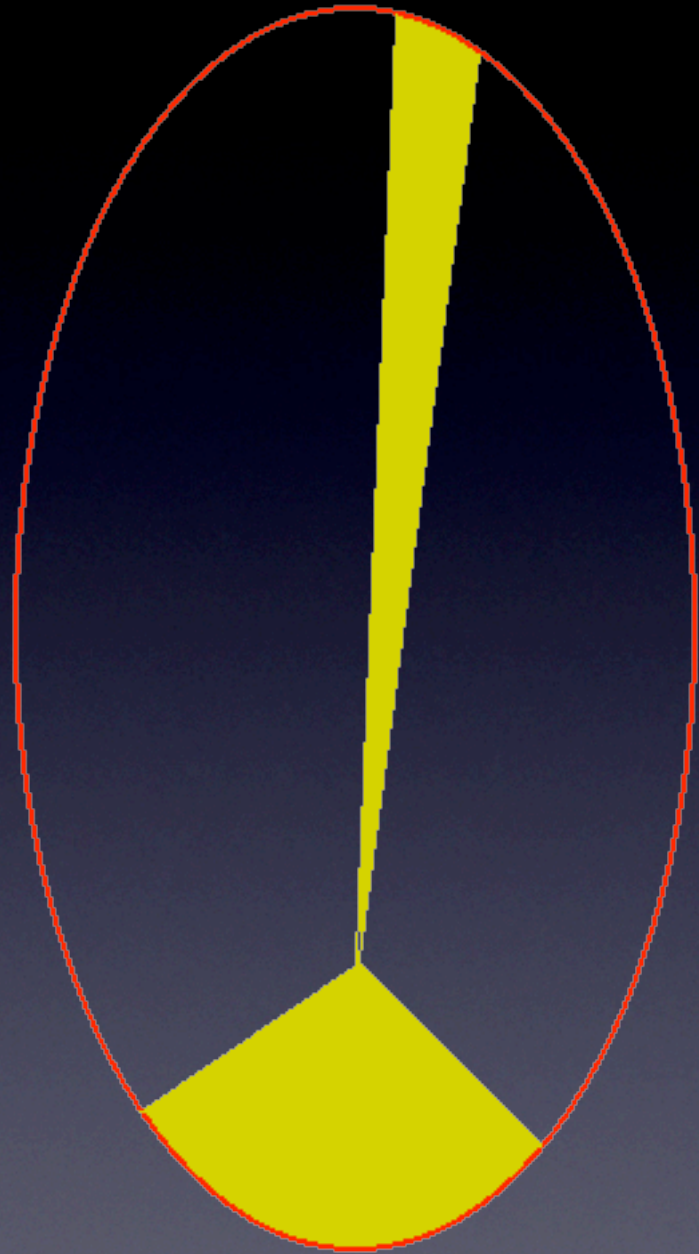
Solis orbis & Stellarium fixarum sphaera immobilis

Mercurius octuaginta diebus Luna mensis Venus nono mense Telluris Luna mensis Martis bima reuolutio Iouis. XII. annorum reuolutio Saturnus anno. XXX. reuoluitur

Kepler

ellipses!

Kepler's laws of planetary motion
better for ephemerides
simpler



Newton

calculus

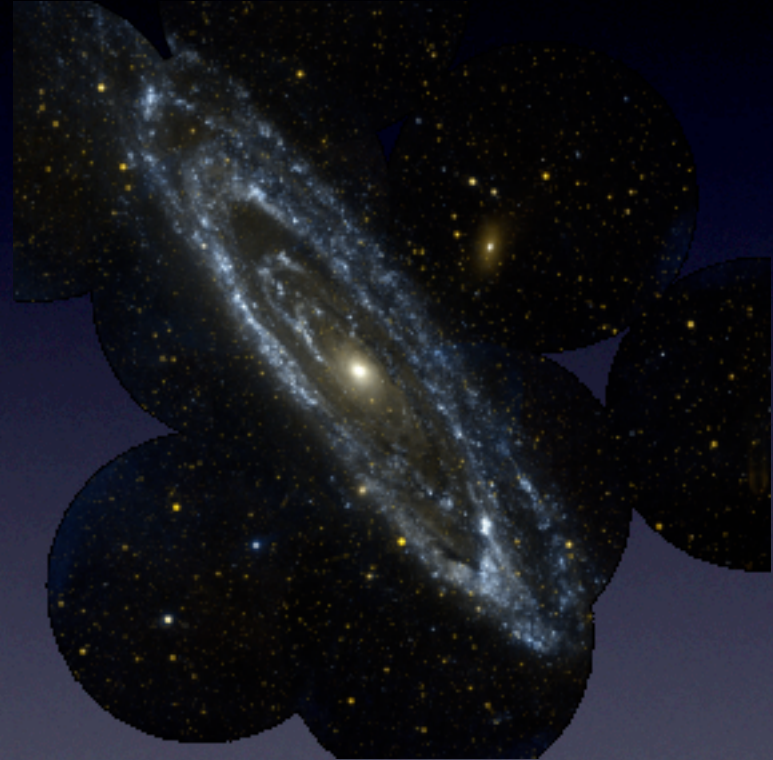
The Law of Universal Gravitation
as above, so below
celestial mechanics

$$f = m \frac{dv}{dr}$$

$$f = \frac{Gmm'}{r^2}$$

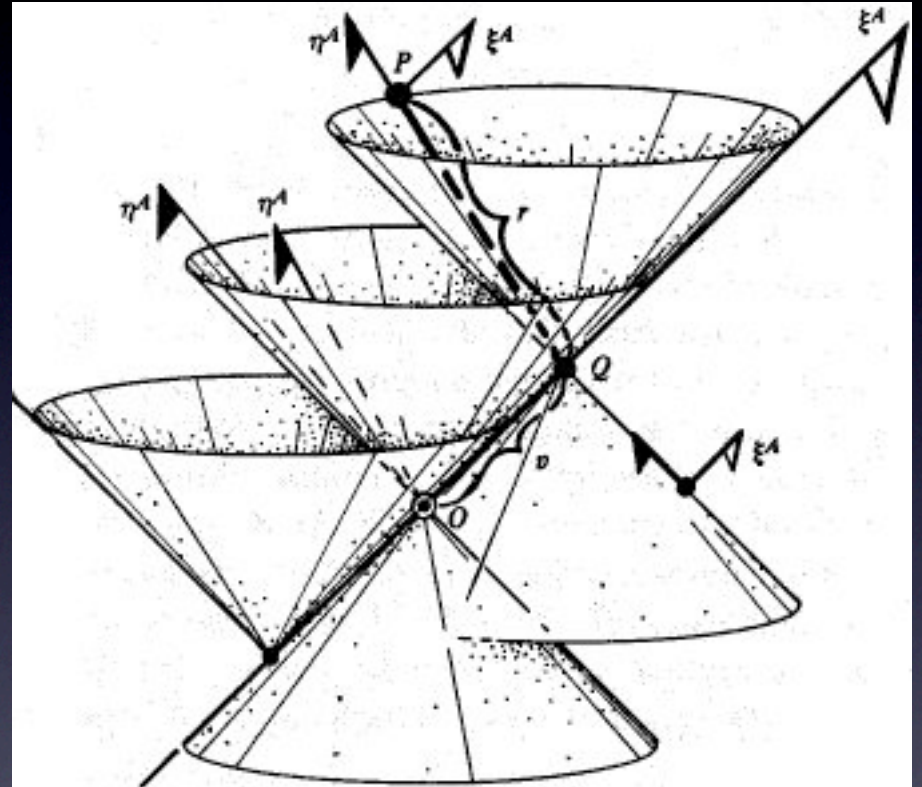
Hubble

distance to Andromeda
too far away to be in our galaxy
...but in our local group
...remoter galaxies are receding

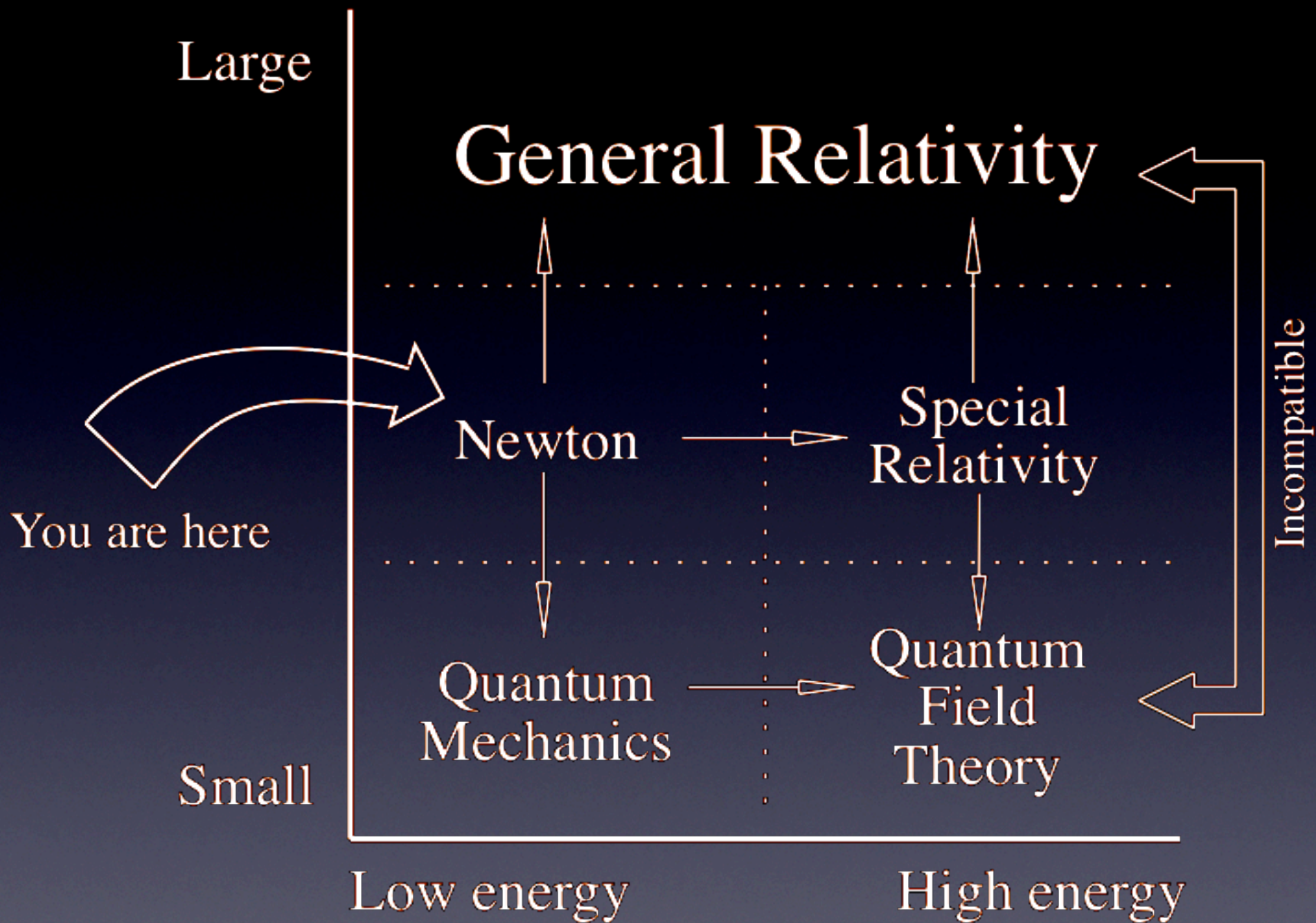


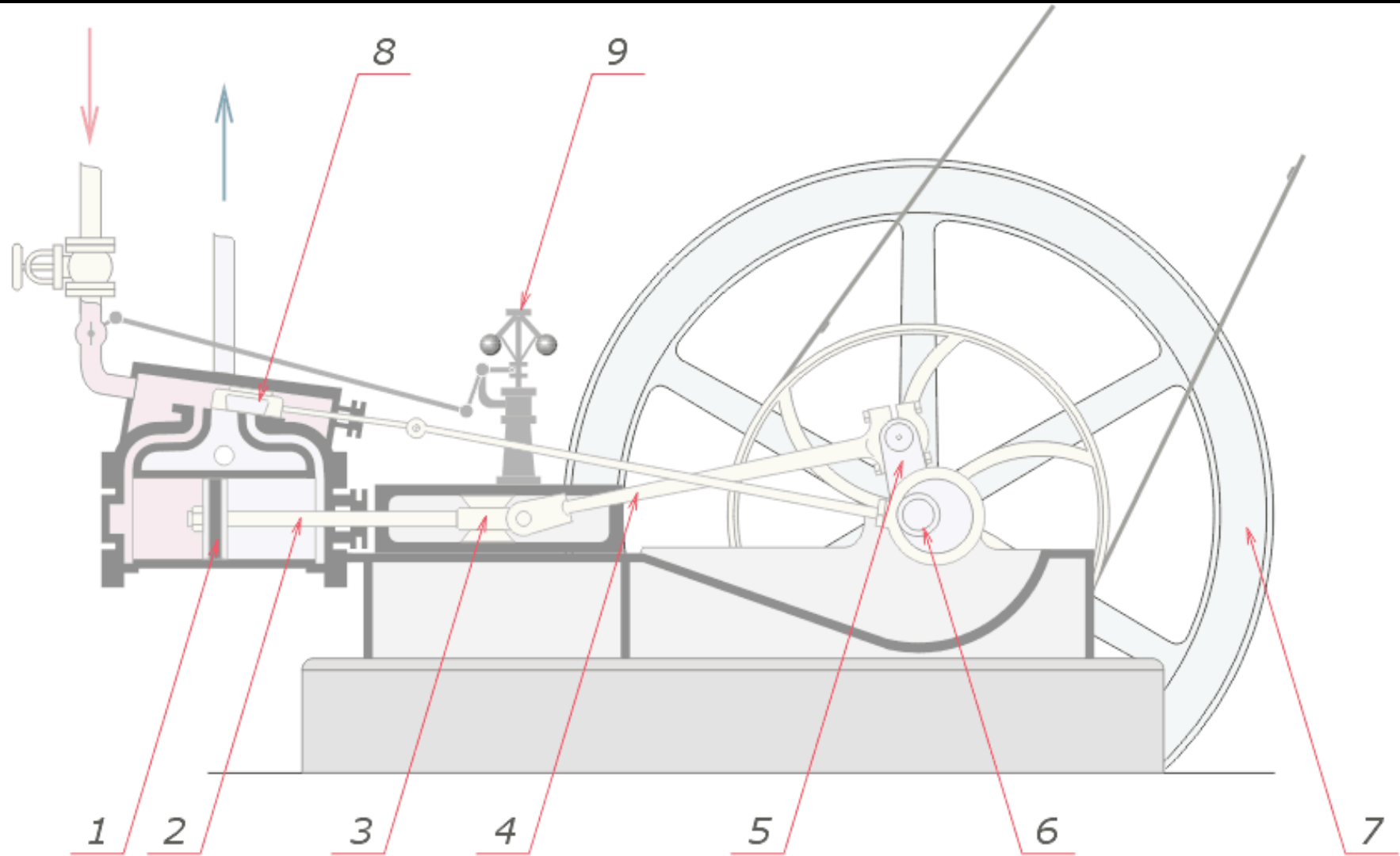
Einstein

geometry of spacetime
large-scale structure of spacetime
highly mathematical
...but still fundamentally simple!



The Standard Model





$$S = Ak_{BC}^3 / 4hG$$

The particle zoo

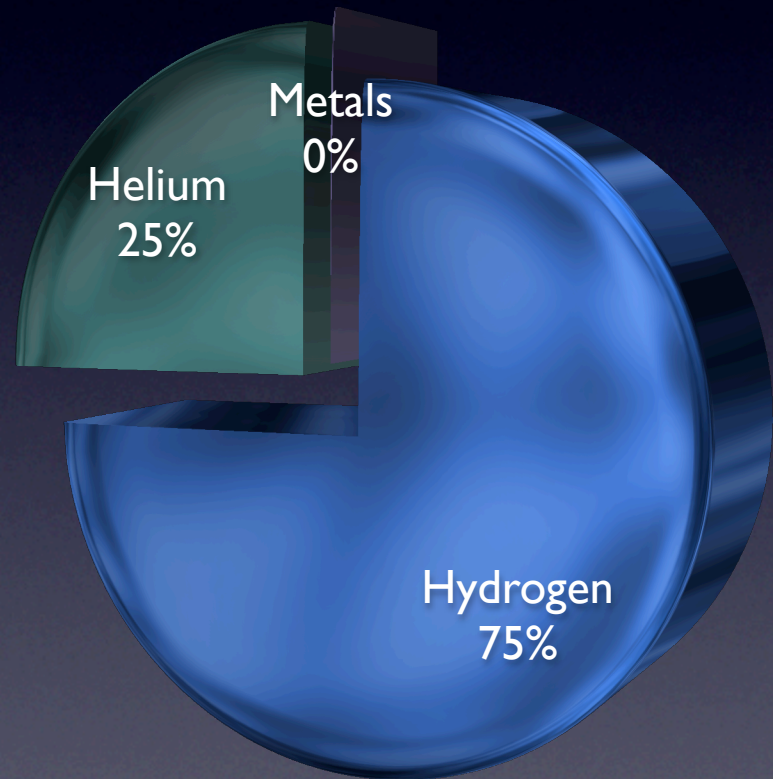
- **Electromagnetism** – This is the force that drives motors, comprises radio waves, keeps the electrons circling the nucleus, and stops us falling through the floor.
- **Weak nuclear force** – Responsible for radioactive decay, when a neutron decays into a proton, and electron, and a neutrino.
- **Strong nuclear force** – The strong nuclear force is the force that keeps nuclei firmly in one piece against the huge electrostatic repulsive force.

The particle zoo (2)

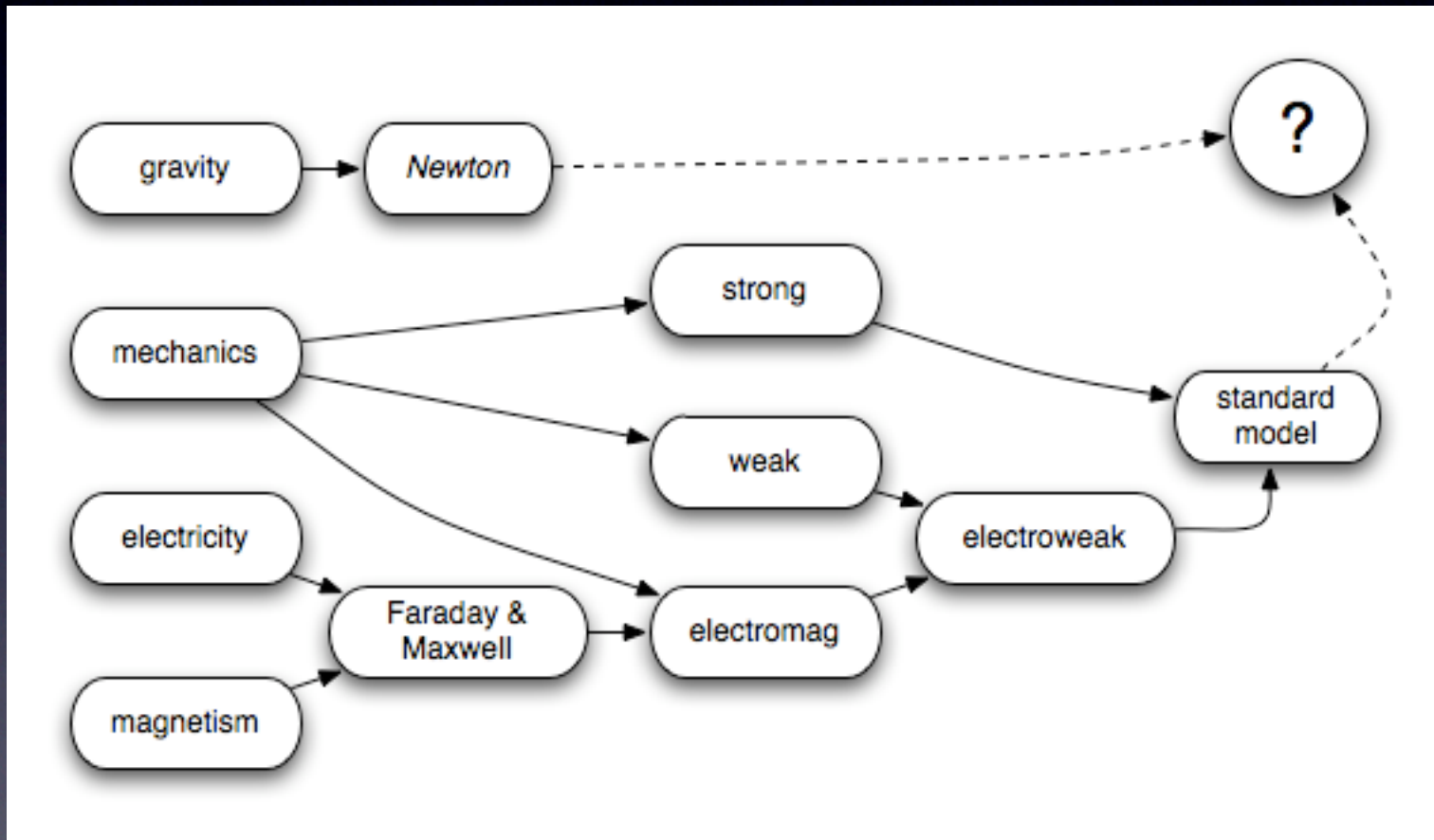
- **Gravity** – In this picture, gravity is the stage on which all of the rest of physics takes place, and in consequence *everything* feels its effects, even the massless photon. ‘Space tells matter how to move; matter tells space how to curve.’

Big bang nucleosynthesis

alpher-bethe-gamow paper



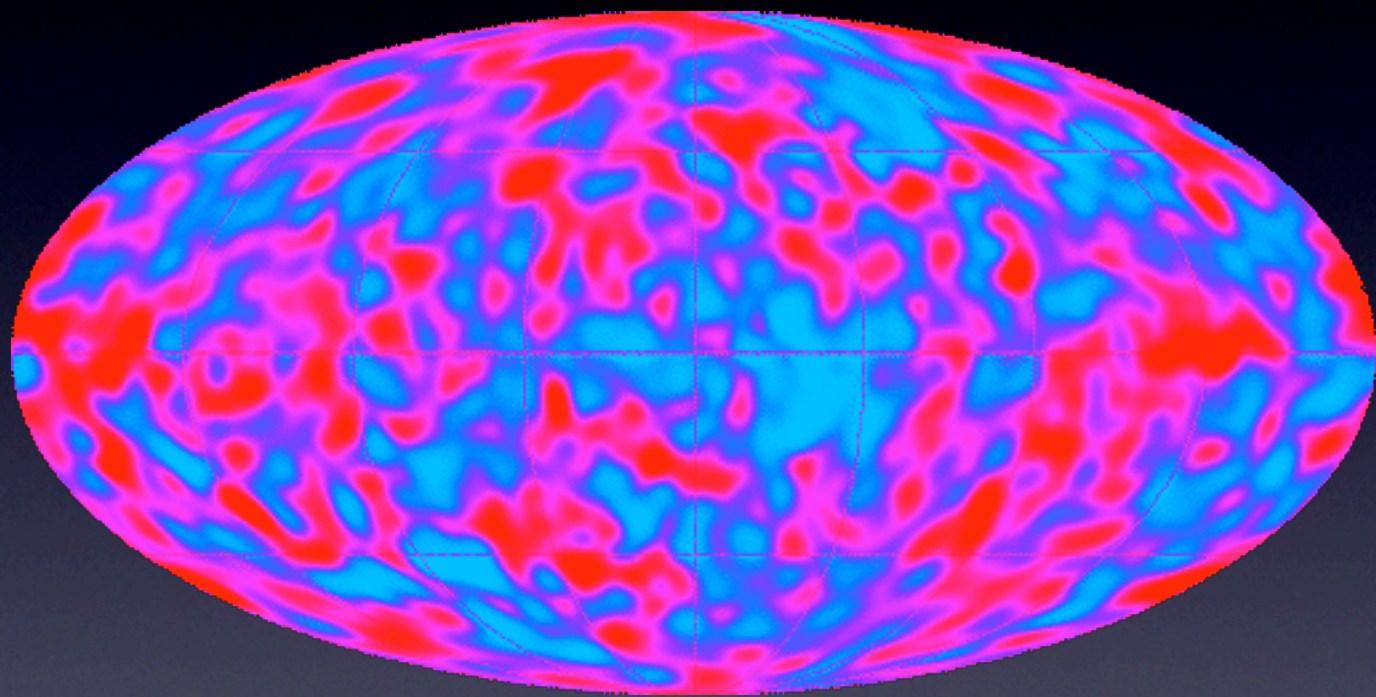
Unification

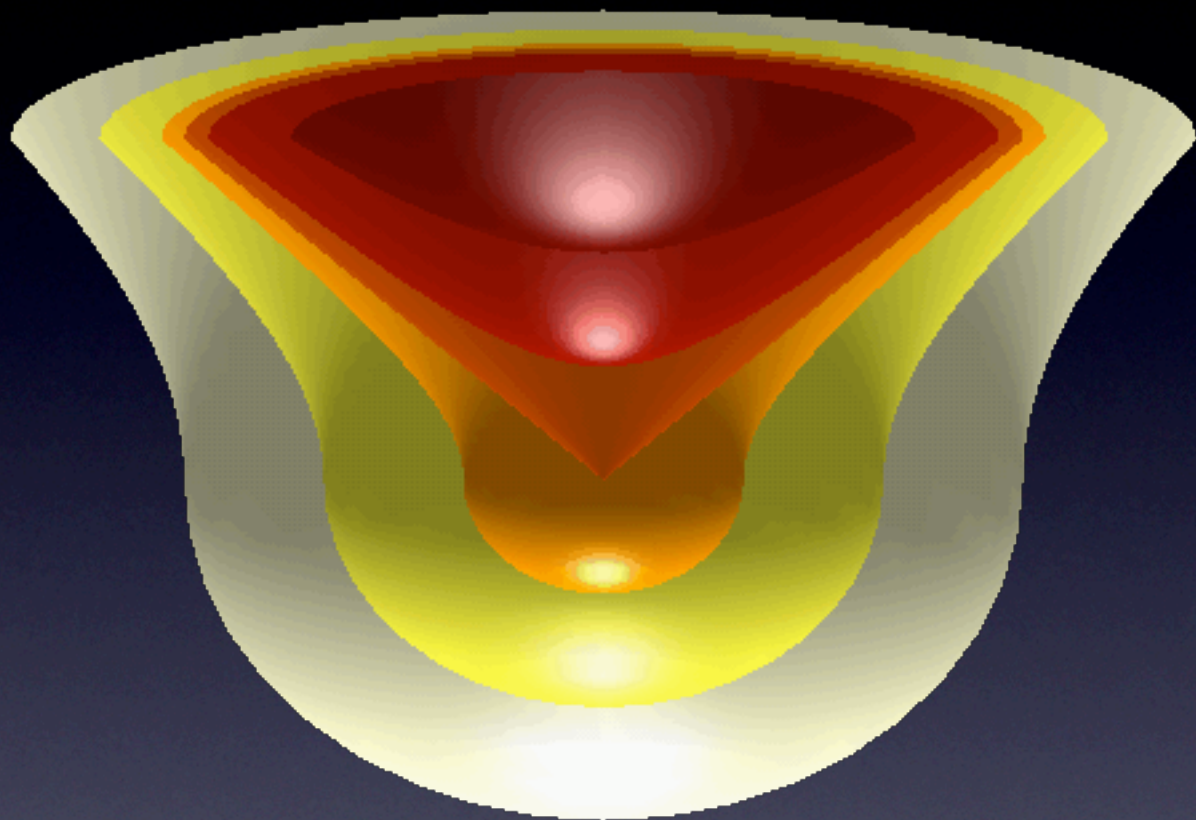


Cosmogeny

Big bang timeline

- $10^{-37}\text{s} - 10^{-35}\text{s}$: inflation
- $10^{-35}\text{s} - 1\text{s}$: particle physics
- $1\text{s} - 3\text{min}$: nuclear physics, nucleosynthesis
- $3\text{min} - 300\,000\text{y}$: radiation, to the CMBR
- 100My : the first stars
- $100\text{My} - 10^{10}\text{y}$: stars, planets, people, stuff





What drives inflation?

Symmetry breaking of an underlying field
...which is therefore an actor
73% dark energy
23% dark matter
4% matter

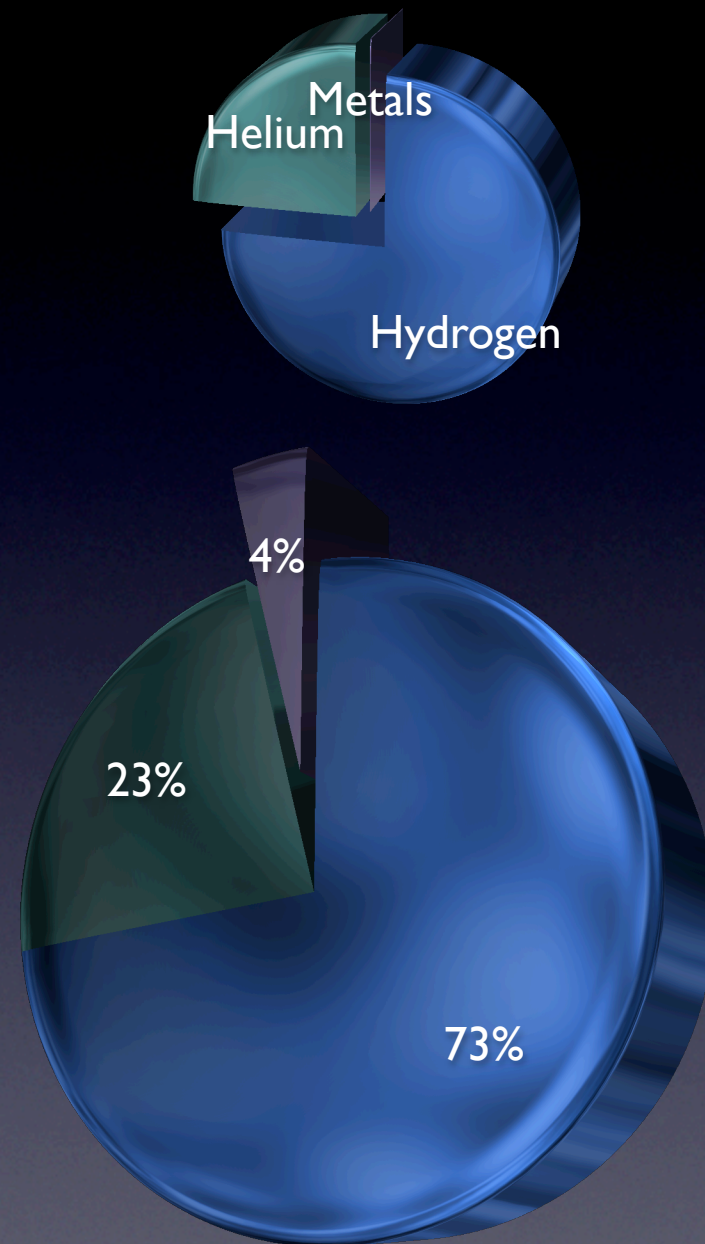


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