2. More Theoretical Foundations: Marginalisation

Suppose there are a set of M propositions $\{x_k : k = 1, ..., M\}$

Then
$$\sum_{k=1}^{M} p(x_k \mid I) = 1$$





2. More Theoretical Foundations: Marginalisation

Suppose there are a set of M propositions $\{x_k : k = 1, ..., M\}$

Then
$$\sum_{k=1}^{M} p(x_k \mid I) = 1$$

Suppose we introduce some additional proposition Y

Use Bayes' theorem.

$$p(x_1, y | I) = p(x_1 | y, I)p(y | I)$$

$$\vdots$$

$$p(x_M, y | I) = p(x_M | y, I)p(y | I)$$



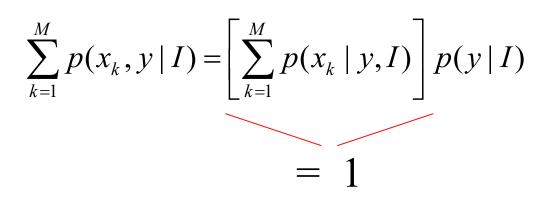


$$\sum_{k=1}^{M} p(x_k, y \mid I) = \left[\sum_{k=1}^{M} p(x_k \mid y, I)\right] p(y \mid I)$$
$$= 1$$









Marginal probability

$$p(y \mid I) = \sum_{k=1}^{M} p(x_k, y \mid I)$$





This extends to the continuum limit :

x can take infinitely many values

$$p(y | I) = \int_{-\infty}^{\infty} p(x, y | I) dx$$





This extends to the continuum limit :

x can take infinitely many values

$$p(y \mid I) = \int_{-\infty}^{\infty} p(x, y \mid I) \, dx$$

p(x, y | I) is no longer a probability, but a probability density

Prob
$$(a \le x \le b \text{ and } y \text{ is true } | I) = \int_{a}^{b} p(x, y | I) dx$$

with obvious extension to continuum limit for y





This extends to the continuum limit :

x can take infinitely many values

$$p(y \mid I) = \int_{-\infty}^{\infty} p(x, y \mid I) \, dx$$

Also
$$\int_{-\infty}^{\infty} p(x \mid y, I) dx = 1$$
Normalisation condition



SUPA)

Probabilities are never negative, so $p(x) \ge 0$ for all x

We compute probabilities by measuring the area under the pdf curve, i.e.

Prob
$$(a \le x \le b) = \int_{a}^{b} p(x) dx$$

'Normalisation' $\int_{-\infty}^{\infty} p(x) dx = 1$





X

b

а

1) Poisson pdf

e.g. number of photons / second counted by a detector, number of objects counted by a survey

r = number of detections

Poisson pdf assumes detections are independent, and there is a constant rate $\ \mu$





1) Poisson pdf

e.g. number of photons / second counted by a detector, number of objects counted by a survey

r = number of detections

$$p(r) = \frac{\mu^r e^{-\mu}}{r!}$$

Poisson pdf assumes detections are independent, and there is a constant rate $\ \mu$



1) Poisson pdf

e.g. number of photons / second counted by a detector, number of objects counted by a survey

r = number of detections

$$p(r) = \frac{\mu^r e^{-\mu}}{r!}$$

Poisson pdf assumes detections are independent, and there is a constant rate $\ \mu$

Can show that

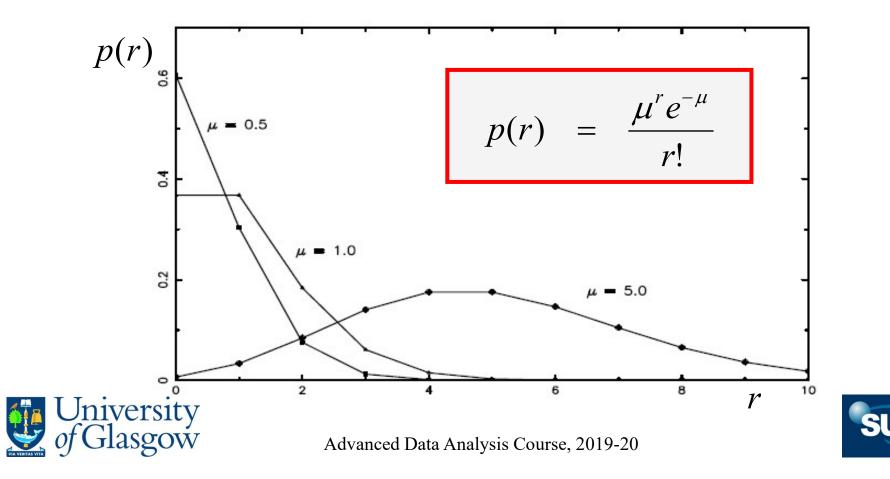
$$\sum_{r=0}^{\infty} p(r) =$$





1) Poisson pdf

e.g. number of photons / second counted by a detector, number of objects counted by a survey



2) Binomial pdf

number of 'successes' from $N\,$ observations, for two mutually exclusive outcomes (e.g. 'Heads' and 'Tails')

r = number of 'successes'

 θ = probability of 'success' for single observation

$$p_N(r) = \frac{N!}{r!(N-r)!}\theta^r(1-\theta)^{N-r}$$





2) Binomial pdf

number of 'successes' from $N\,$ observations, for two mutually exclusive outcomes (e.g. 'Heads' and 'Tails')

r = number of 'successes'

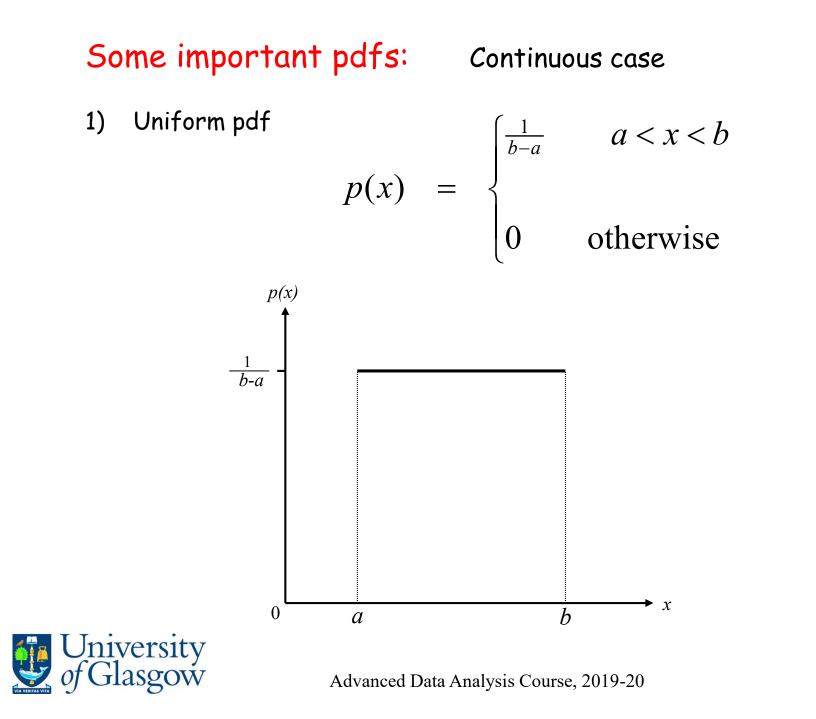
 θ = probability of 'success' for single observation

$$p_N(r) = \frac{N!}{r!(N-r)!} \theta^r (1-\theta)^{N-r}$$

Can show that
$$\sum_{r=0}^{\infty} p_N(r) = 1$$





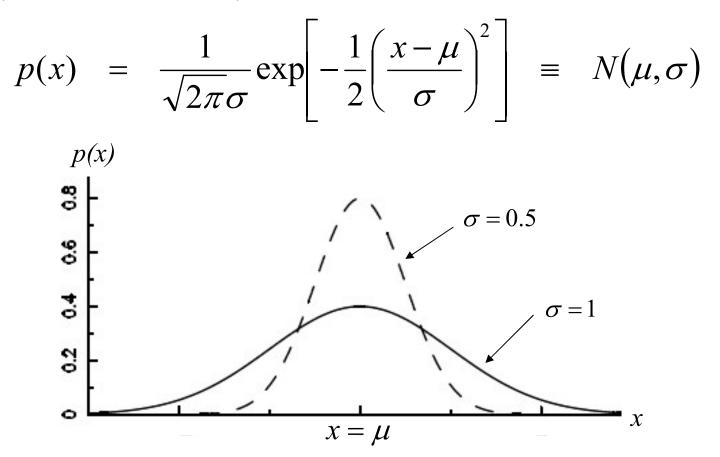




Some important pdfs:

Continuous case

2) Central, or normal pdf (also known as Gaussian)

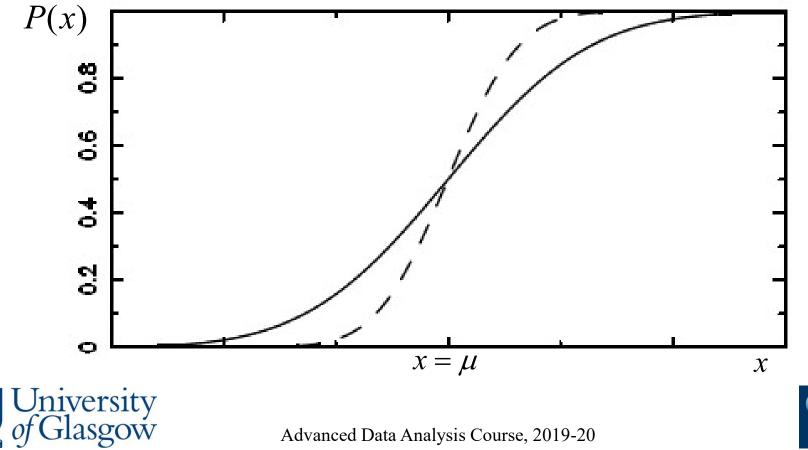






Cumulative distribution function (CDF)

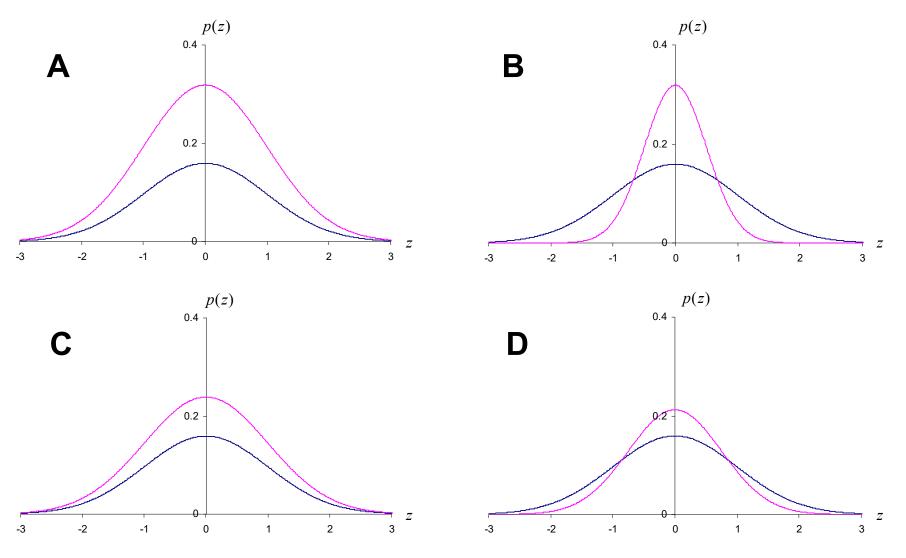
$$P(a) = \int_{-\infty}^{a} p(x) dx = \operatorname{Prob}(x < a)$$





Question 1: In the figures below, the blue curves show a normal distribution with mean zero and $\sigma = 1$.

Which of the pink curves shows a normal distribution with mean zero and $\sigma = 0.5$?



The nth moment of a pdf is defined as:-

$$\langle x^n \rangle = \sum_{x=0}^{\infty} x^n p(x | I)$$
 Discrete case
 $\langle x^n \rangle = \int_{-\infty}^{\infty} x^n p(x | I) dx$ Continuous case





The 1st moment is called the **mean** or **expectation value**:

$$E(x) = \langle x \rangle = \sum_{x=0}^{\infty} x p(x | I)$$
 Discrete case
$$E(x) = \langle x \rangle = \int_{-\infty}^{\infty} x p(x | I) dx$$
 Continuous case



The 2nd moment is called the **mean square**:

$$\left\langle x^{2} \right\rangle = \sum_{x=0}^{\infty} x^{2} p(x | I)$$
 Discrete case
 $\left\langle x^{2} \right\rangle = \int_{-\infty}^{\infty} x^{2} p(x | I) dx$ Continuous case





The variance is defined as:

$$\operatorname{var}[x] = \sum_{x=0}^{\infty} (x - \langle x \rangle)^2 p(x \mid I) \qquad \text{Discrete case}$$
$$\operatorname{var}[x] = \int_{-\infty}^{\infty} (x - \langle x \rangle)^2 p(x \mid I) dx \qquad \text{Continuous case}$$

and is often written as σ^2

 $\sigma = \sqrt{\sigma^2}$ is called the standard deviation





The variance is defined as:

$$\operatorname{var}[x] = \sum_{x=0}^{\infty} (x - \langle x \rangle)^2 p(x \mid I) \qquad \text{Discrete case}$$
$$\operatorname{var}[x] = \int_{-\infty}^{\infty} (x - \langle x \rangle)^2 p(x \mid I) dx \qquad \text{Continuous case}$$

In general
$$\operatorname{var}[x] = \langle x^2 \rangle - \langle x \rangle^2$$





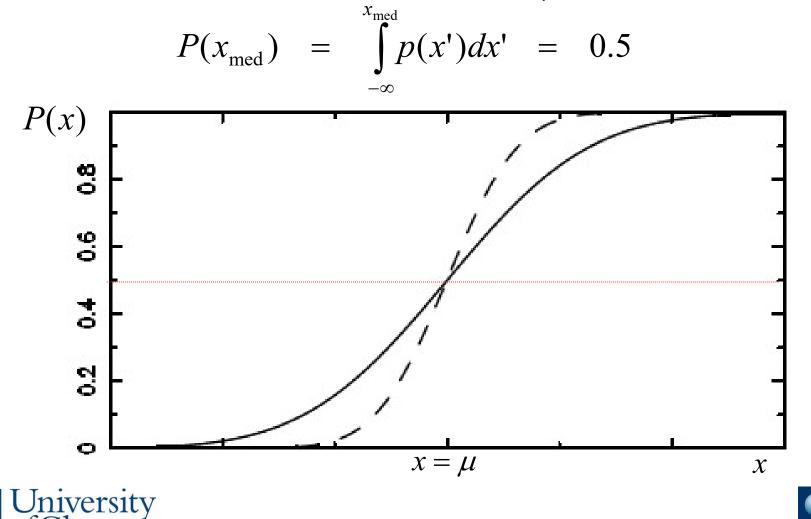
pdf	mean	variance
Poisson $p(r) = \frac{\mu^r e^{-\mu}}{r!}$	μ	μ
Binomial $p_N(r) = \frac{N!}{r!(N-r)!} \theta^r (1-\theta)^{N-r}$	N heta	$N\theta(1-\theta)$
Uniform $p(x) = \frac{1}{b-a}$	$\frac{1}{2}(a+b)$	$\frac{1}{12}(b-a)^2$
Normal $\begin{bmatrix} 1 & (1 - 1)^2 \end{bmatrix}$		2
$p(x) = \frac{1}{\sqrt{2\pi\sigma}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^2\right]$	μ	$ \sigma^2$

Note: for a normal pdf all the higher moments are zero





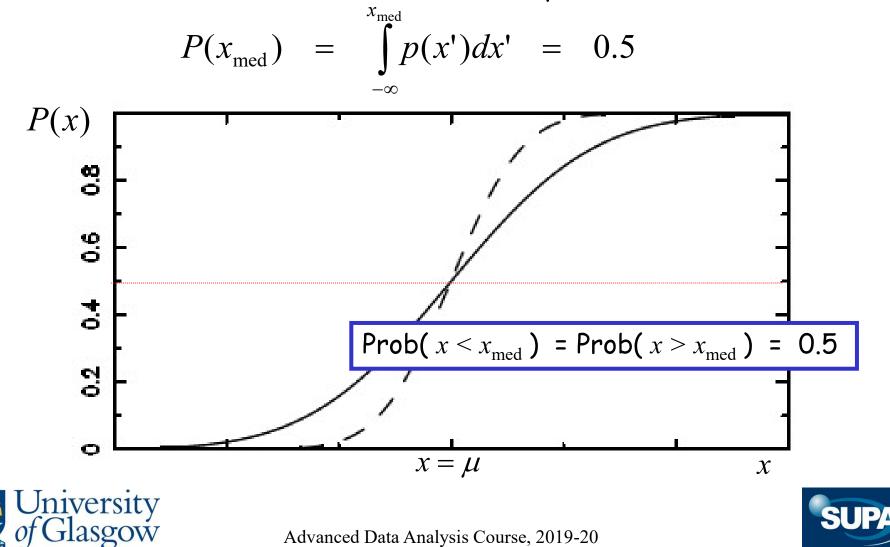
The Median divides the CDF into two equal halves







The Median divides the CDF into two equal halves



Advanced Data Analysis Course, 2019-20

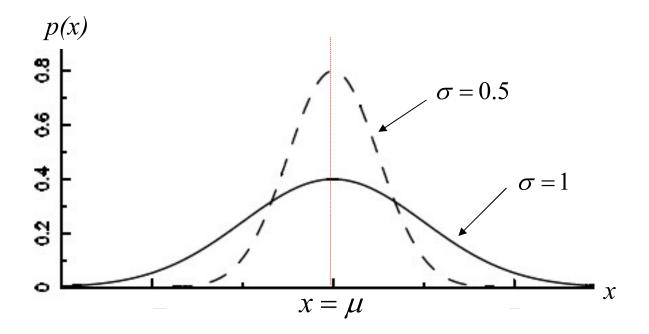
Question 2: The following table lists the absolute magnitude (related to the log of luminosity) of the brightest galaxy in 15 distant galaxy clusters. The galaxies are arranged from brightest to faintest.

Cluster #	M_{ν}	Cluster #	M_{V}	Cluster #	M_{V}
1	-23.36	6	-22.37	11	-21.14
2	-23.23	7	-21.89	12	- 2 <mark>1</mark> .13
3	- 22.50	8	-21.33	13	- <mark>21.1</mark> 1
4	-22.48	9	-21.23	14	-21.08
5	- 22.46	10	-21.17	15	-21.04

The sample median of these data is

A
$$-23.36$$
 B -21.04 **C** -21.33 **D** -21

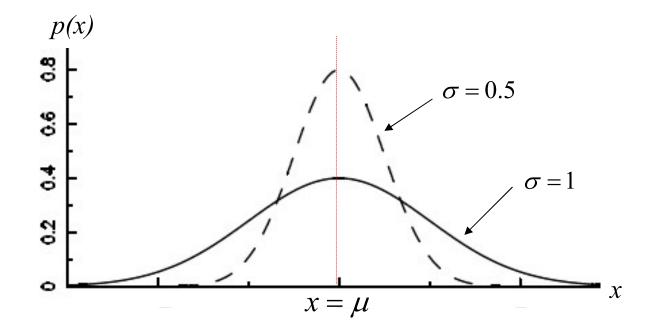
The Mode is the value of x for which the pdf is a maximum







The Mode is the value of x for which the pdf is a maximum



For a normal pdf, mean = median = mode = μ



SUPA)

Variance of a Function of a RV

The variance, var[f(x)], of an arbitrary function of x can be approximated to second order by the following expression

$$\operatorname{var}[f(x)] = \operatorname{var}(x) \left(\frac{\partial f}{\partial x}\right)_{x=\overline{x}}^2$$

This expression is the basis for the 'error propagation' formulae we use in e.g. undergraduate physics labs

See also the SUPAIDA course





Question 3: Which expression correctly approximates the error on the *natural logarithm* of a variable x ?

 $\mathbf{A} \qquad \sigma_{\ln x} \sim \frac{\sigma_x^2}{r^2}$

B $\sigma_{\ln x} \sim x^2 \sigma_x^2$

C $\sigma_{\ln x} \sim \frac{\sigma_x}{\chi}$

D $\sigma_{\ln x} \sim x \sigma_x$

Multivariate Distributions

Thus far we have considered only the pdf of a single (univariate) RV. We now extend to the **multivariate** case of two or more RVs.

Joint pdf

The joint pdf of two RVs, x_1 and x_2 is $p(x_1, x_2)$. Then,

$$\operatorname{Prob}(a_1 < X_1 < b_1 \text{ and } a_2 < X_2 < b_2) = \int_{a_1}^{b_1} \int_{a_2}^{b_2} p(x_1, x_2) \, dx_1 dx_2$$

Extension to more than two RVs is carried out in the obvious way.





Marginal Distributions

The marginal pdf, $p_1(x_1)$ of x_1 is defined by

$$p_1(x_1) = \int_{-\infty}^{\infty} p(x_1, x_2) \, dx_2$$

and is a pdf in the usual sense that

1.
$$p_1(x_1) \ge 0$$
, for all x_1

2.
$$\operatorname{Prob}(a < x_1 < b) = \int_a^b p_1(x_1) dx_1$$

3.
$$\int_{-\infty}^{\infty} p_1(x_1) dx_1 = 1$$





Marginal Distributions

Similarly, the marginal pdf of x_2 is

$$p_2(x_2) = \int_{-\infty}^{\infty} p(x_1, x_2) \, dx_1$$

In general, given any multivariate pdf, we may find the marginal pdf of any subset of the $x_1, ..., x_n$ by integrating over all other variables. e.g.

$$p_{13}(x_1, x_3) = \int_{-\infty}^{\infty} \dots \int_{-\infty}^{\infty} p(x_1, \dots, x_n) \, dx_2 dx_4 dx_5 \dots dx_n$$



SUPA)

Conditional Distributions

Consider the joint pdf, $p(x_1, x_2)$, of x_1 and x_2 . Suppose we observe x_1 , but do not observe x_2 . We want a function that describes the pdf of x_2 , given the observed value of x_1 (usually simply stated as 'given x_1 '). This function is known as the **conditional** pdf of x_2 , written as $p(x_2|x_1)$, and defined by

$$p(x_2|x_1) = \frac{p(x_1, x_2)}{p_1(x_1)}$$

i.e. the conditional pdf is obtained by dividing the joint pdf of x_1 and x_2 by the marginal pdf of x_1 (provided $p_1(x_1) \neq 0$).





Conditional Distributions

Similarly

$$p(x_1|x_2) = \frac{p(x_1, x_2)}{p_2(x_2)}$$

Extension to more than 2 RVs is again straightforward. For exam-

ple,

$$p(x_1, x_3 | x_2, x_4) = \frac{p(x_1, x_2, x_3, x_4)}{p_{24}(x_2, x_4)}$$





Statistical Independence

If the conditional pdf of x_2 given x_1 does not depend on x_1 , this means that x_1 and x_2 are statistically independent, since the observed value of x_2 is unaffected by the observed value of x_1 .

Equivalently, x_1 and x_2 are independent if and only if the joint pdf of x_1 and x_2 can be written as the product of their marginal pdfs, i.e.

$$p(x_1, x_2) = p_1(x_1) p_2(x_2)$$

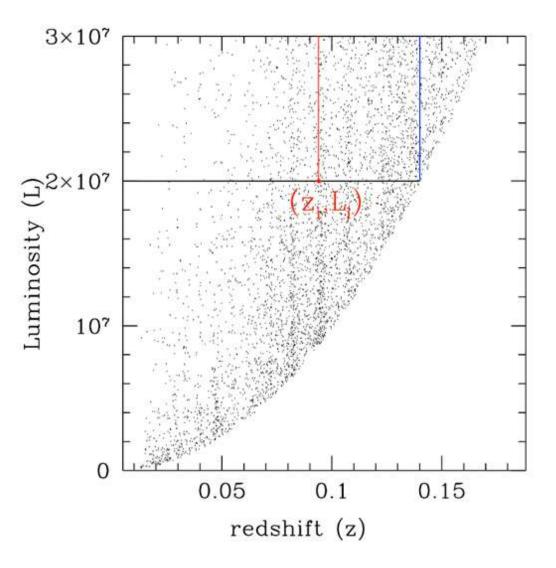


Example application

Is the **luminosity** distribution of galaxies independent of **redshift**, or do galaxies evolve?

To answer this question is not straightforward. The *observed* distributions of luminosity and redshift are in any case not independent – due to what are known as **selection effects**.

At higher redshift only the more luminous galaxies are bright enough to be detected, so the observed luminosity distribution *appears* to evolve – but is this evolution real?...







Question 4: Which of the following joint pdfs describe variables x and y which are statistically independent?

A
$$p(x, y) \propto \frac{1}{2}(x+y);$$
 $0 \le x, y < \infty$

B
$$p(x, y) \propto \exp\left[-\frac{1}{2}(x+y)\right]; \quad 0 \le x, y < \infty$$

C
$$p(x, y) \propto \log(x + y); \quad 0 < x, y < \infty$$

D
$$p(x, y) \propto \exp\left[-\frac{1}{2}(x+y)\right]; \quad 0 \le x < y, \ 0 \le y < \infty$$

Let x and y be RVs with the following joint pdf

$$p(x,y) = \frac{1}{2\pi\sigma_{\rm x}\sigma_{\rm y}\sqrt{1-\rho^2}} \exp\left[-\frac{1}{2(1-\rho^2)}Q(x,y)\right]$$

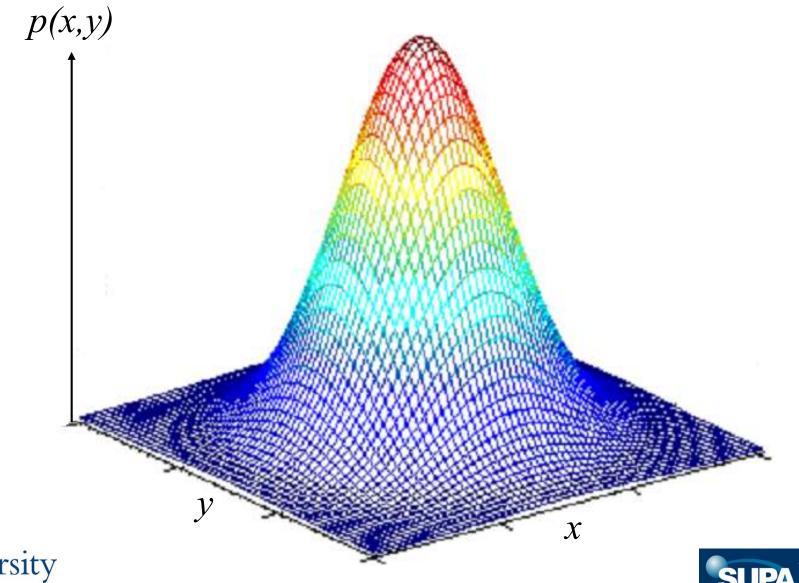
where the quadratic form, Q(x, y) is given by

$$Q(x,y) = \left(\frac{x-\mu_{\mathbf{x}}}{\sigma_{\mathbf{x}}}\right)^2 - 2\rho\left(\frac{x-\mu_{\mathbf{x}}}{\sigma_{\mathbf{x}}}\right)\left(\frac{y-\mu_{\mathbf{y}}}{\sigma_{\mathbf{y}}}\right) + \left(\frac{y-\mu_{\mathbf{y}}}{\sigma_{\mathbf{y}}}\right)^2$$

Then p(x, y) is known as the **bivariate normal pdf** and is specified by the 5 parameters μ_x , μ_y , σ_x , σ_y and ρ . This pdf is used often in the physical sciences to model the joint pdf of two random variables.











The first 4 parameters of the bivariate normal pdf are, in fact, equal to the following expectation values:-

1.
$$E(x) = \mu_x$$

- 2. $E(y) = \mu_y$
- 3. $\operatorname{var}(x) = \sigma_{\mathbf{x}}^2$
- 4. $\operatorname{var}(y) = \sigma_y^2$



The parameter ρ is known as the **correlation coefficient** and satisfies

$$E[(x - \mu_{\rm x})(y - \mu_{\rm y})] = \rho \sigma_{\rm x} \sigma_{\rm y}$$

Note that if $\rho = 0$ then x and y are independent.

 $E[(x - \mu_x)(y - \mu_y)]$ is known as the **covariance** of x and y and is often denoted by cov(x, y).





The parameter ρ is known as the **correlation coefficient** and satisfies

$$E[(x - \mu_{\rm x})(y - \mu_{\rm y})] = \rho \sigma_{\rm x} \sigma_{\rm y}$$

Note that if $\rho = 0$ then x and y are independent.

 $E[(x - \mu_x)(y - \mu_y)]$ is known as the **covariance** of x and y and is often denoted by cov(x, y).

In fact, for any two variables x and y, we define

$$\operatorname{cov}(x, y) = E[(x - E(x))(y - E(y))]$$

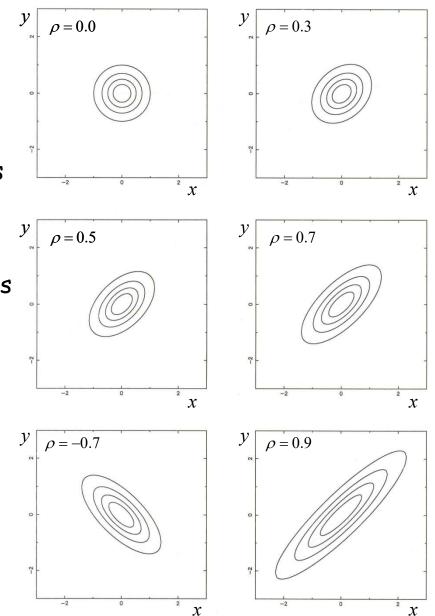




Isoprobability contours for the bivariate normal pdf

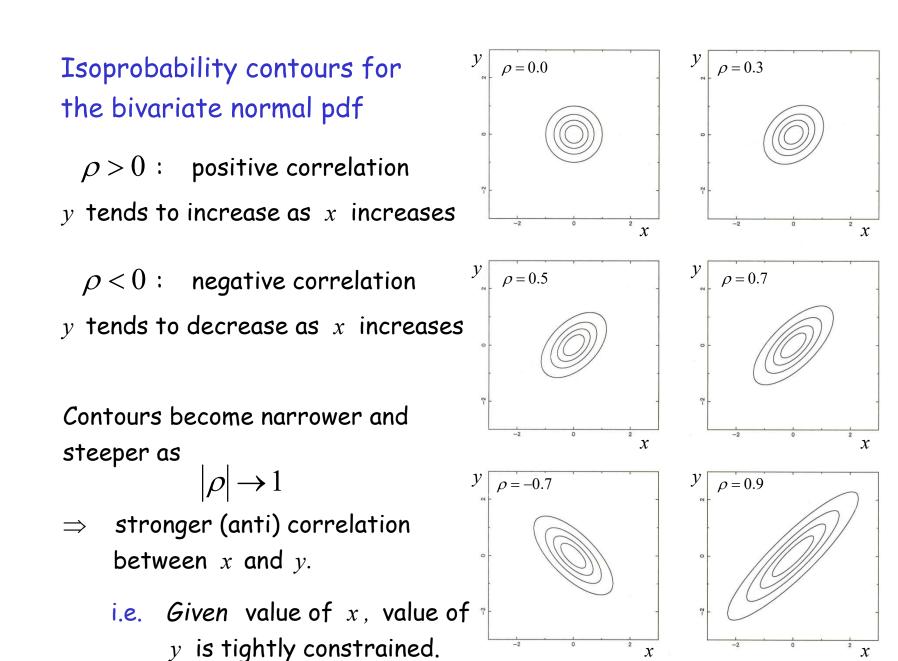
 $\rho > 0$: positive correlation y tends to increase as x increases

 $\rho < 0$: negative correlation y tends to decrease as x increases









urse 2019-20





The marginal pdfs of x and y are just the univariate normal pdfs, i.e.

$$p_x(x) = N(\mu_x, \sigma_x)$$
 $p_y(y) = N(\mu_y, \sigma_y)$

The conditional pdf of y given x is also a univariate normal pdf, viz:-

$$p(y|x) = N(\mu_{y} + \frac{\sigma_{y}}{\sigma_{x}}\rho(x - \mu_{x}), \sigma_{y}\sqrt{1 - \rho^{2}})$$

with the corresponding expression for p(x|y).





 $\mu_{y} + \frac{\sigma_{y}}{\sigma_{x}}\rho(x - \mu_{x})$ is often referred to as the **conditional expectation** (value) of y given x, and the equation

$$y = \mu_{\rm y} + \frac{\sigma_{\rm y}}{\sigma_{\rm x}}\rho(x - \mu_{\rm x})$$

is called the **regression line** of y on x.



