

HEAT II Ex Sheet 3Model Answers

1) Assuming the source is isotropic, $L_x = 4\pi d^2 F$

$$d = 50 \times 3.09 \times 10^{19} \text{ m} \quad 5 \text{ keV} = 8 \times 10^{-16} \text{ J}$$

$$F = \left(\frac{50 \times 8 \times 10^{-16}}{0.04} \right) \text{ Wm}^{-2} = 10^{-12} \text{ Wm}^{-2}$$

$$\begin{aligned} \Rightarrow L_x &= 4\pi \times (1.545 \times 10^{21})^2 \times 10^{-12} \text{ W} \\ &= 3 \times 10^{31} \text{ W} \end{aligned}$$

$$2) L_x = \int_{0.5}^{10} \frac{dL}{d\varepsilon} d\varepsilon = \int_{0.5}^{10} 2 \left(\frac{2}{\pi m_e} \right)^{1/2} \frac{n_p^2 V Q_0 m_e c^2}{(kT)^{1/2}} e^{-\varepsilon/kT} d\varepsilon$$

Our goal is to determine n_p

$$kT = 1.38 \times 10^{-15} \text{ J} = 8.6 \text{ keV}, \text{ so putting } y = \varepsilon/kT$$

$$\text{when } \varepsilon = 0.5, \quad y = 0.058 = y_1$$

$$\varepsilon = 10, \quad y = 1.163 = y_2$$

$$L_x = 2 \left(\frac{2}{\pi m_e} \right)^{1/2} (kT)^{1/2} V Q_0 m_e c^2 \int_{y_1}^{y_2} e^{-y} dy n_p^2$$

$$= \text{const.} \times \left[-e^{-y} \right]_{y_1}^{y_2} n_p^2$$

$$\text{const.} = 2 \times \left(\frac{2}{\pi \times 9.1 \times 10^{-31}} \right)^{1/2} \times (1.38 \times 10^{-15})^{1/2} \times 2.67 \times 10^{67} \times 1.54 \times 10^{-31} \times 9.1 \times 10^{-31} \times (3 \times 10^8)^2$$

$$= 2 \times 10^{31} \text{ (in SI units)}$$

(2)

$$\text{So, } 5 \times 10^{37} = 2 \times 10^{31} \times \left[e^{-0.058} - e^{-1.163} \right] n_p^2$$
$$\Leftrightarrow n_p = 2 \times 10^3 \text{ protons / m}^3$$

$$\text{Total mass of intracluster gas} = m_p n_p V$$
$$= 8.9 \times 10^{43} \text{ kg}$$

$$\text{Thermal energy} = 3 N_p k T$$
$$= 3 n_p V k T$$
$$= 3 \times 2 \times 10^3 \times 2.67 \times 10^{67} \times 1.38 \times 10^{-15} \text{ J}$$
$$= 2.2 \times 10^{56} \text{ J}$$

$$\text{Cooling time } \tau \sim E_L = \frac{2.2 \times 10^{56}}{5 \times 10^{37}} = 4.4 \times 10^{18} \text{ s}$$
$$= 1.4 \times 10^{11} \text{ yrs}$$

3) Assuming $U_B \sim U_{\text{opt}} = 10^6 \text{ eV m}^{-3}$

$$= 1.6 \times 10^{-13} \text{ J m}^{-3}$$
$$= \frac{B^2}{2\mu_0}$$

$$\Rightarrow B^2 = 1.6 \times 10^{-13} \times 2 \times 1.26 \times 10^{-6}$$
$$B = 6.4 \times 10^{-10} \text{ T}$$

A photon of $\lambda = 600 \text{ nm}$ has $\nu = 5 \times 10^{14} \text{ Hz} \Rightarrow E_\nu = 3.3 \times 10^{-19} \text{ J}$

$$\therefore U_{\text{opt}} \text{ corresponds to } \left(1.6 \times 10^{-13} / 3.3 \times 10^{-19} \right) = 4.8 \times 10^5 \text{ photons m}^{-3}$$

$$\text{Taking } \nu_L = \frac{eB}{2\pi m_e} = \frac{1.6 \times 10^{-19} \times 6.4 \times 10^{-10}}{2\pi \times 9.1 \times 10^{-31}} = 17.9 \text{ Hz}$$

(3)

$$\text{We want } \nu_s = 10^{18} \text{ Hz} \Rightarrow \frac{3}{2} \gamma^2 = \frac{10^{18}}{17.9}$$

$$\Leftrightarrow \gamma = 1.9 \times 10^8$$

$$4) \text{ Number density of CMBR photons} = \frac{5 \times 10^{-14}}{3 \times 10^{-4} \times 1.6 \times 10^{-19}} \text{ m}^{-3} \approx 10^9 \text{ photons m}^{-3}$$

$$\Rightarrow n_\gamma / n_H \approx 10^3$$

$$5) \text{ Column density} = 2 \times 10^{26} \text{ H atoms m}^{-2}$$

$$\text{We take this as } 2 \times 10^{26} \text{ H m}^{-2} = n_H \times d \\ = \left(\frac{3 \times 10^{-21}}{1.67 \times 10^{-27}} \right) d$$

$$\Rightarrow d = \left(\frac{1.67 \times 10^{-27} \times 2 \times 10^{26}}{3 \times 10^{-21}} \right) \text{ m}$$

$$= 1.1 \times 10^{20} \text{ m} \approx 3.6 \text{ kpc}$$

If the column density were 2×10^{28} , this would imply a distance of 360 kpc which is too large for our galaxy.

6) Using the thermal bremsstrahlung formula:

$$L_X = 2 \left(\frac{2}{\pi m_e} \right)^{1/2} (kT)^{1/2} V Q_0 m_e c^2 \left[e^{-y_1} - e^{-y_2} \right] n_p^2$$

$$\text{where } y_i = \epsilon_i / kT$$

If we assume that we integrate over X-ray energies from 0.1 keV to 100 keV (i.e. soft to hard)

(4)

$$kT = 1.38 \times 10^{-16} \text{ J} = 0.86 \text{ keV} \Rightarrow y_1 = 0.116, y_2 = 116$$

$$\text{Integrated flux} = 10^{-4} \text{ W m}^{-2}$$

$$\Rightarrow L_x = 4\pi d^2 F \quad (\text{assuming isotropic emission})$$
$$= 2.8 \times 10^{19} \text{ W}$$

Putting all this together, with $V = \pi r^2 h = 9.8 \times 10^{22} \text{ m}^3$

$$L_x = 2 \times \left(\frac{2}{\pi \times 9.1 \times 10^{-31}} \right)^{1/2} (1.38 \times 10^{-16})^{1/2} \times 9.8 \times 10^{22} \times 1.54 \times 10^{-31} \times 9.1 \times 10^{-31} \times 9 \times 10^{16} \times 0.89 \times n_p^2$$

$$\Leftrightarrow n_p^2 = \frac{2.8 \times 10^{19}}{2.2 \times 10^{-14}} = 3.4 \times 10^{16} \text{ m}^{-3}$$

Assuming hydrogen plasma, $n_p = n_e \Rightarrow n_e = 3.4 \times 10^{16} \text{ m}^{-3}$

$$\text{Thermal energy density} = 3n_e kT = 14 \text{ J m}^{-3}$$

$$\text{mag. energy density } U_B = \frac{B^2}{2\mu_0} = \frac{0.2^2}{2 \times 1.26 \times 10^{-6}} = 1.6 \times 10^4 \text{ J m}^{-3}$$

7) Applying the same formulae as in Q.6; $y_1 = 0.06, y_2 = 58.8$

$$L_x = 2 \times \left(\frac{2}{\pi \times 9.1 \times 10^{-31}} \right)^{1/2} (2.76 \times 10^{-16})^{1/2} \times \frac{4}{3} \pi (10^8)^3 \times 1.54 \times 10^{-31} \times 9.1 \times 10^{-31} \times 0.94 \times 10^{28}$$
$$= 1.4 \times 10^{16} \text{ W}$$

$$\text{Mass of plasma} = 10^{14} \times 1.67 \times 10^{-27} \times \frac{4}{3} \pi (10^8)^3 = 7 \times 10^{11} \text{ kg}$$

(5)

$$\text{Mass of each sphere} = 1.75 \times 10^{11} \text{ kg}$$

$$\text{no. of protons in each sphere} = 1.05 \times 10^{38}$$

$$\text{number density} = 1.05 \times 10^{38} / \left(\frac{4}{3} \pi 10^{21} \right) = 2.5 \times 10^{16} \text{ m}^{-3}$$

$$\Rightarrow L_x = 4 \times 2 \times \left(\frac{2}{\pi \times 9.1 \times 10^{-31}} \right)^{\frac{1}{2}} (2.76 \times 10^{-16})^{\frac{1}{2}} \times \frac{4}{3} \pi (10^7)^3 \times 1.54 \times 10^{-31} \times 9.1 \times 10^{-31} \\ \times 9 \times 10^{16} \times 0.94 \times (2.5 \times 10^{16})^2$$

So the volume \downarrow 250, no. density \uparrow $(250)^2$

\Rightarrow luminosity increased by a factor of 250

$$\Rightarrow L_x \approx 3.5 \times 10^{18} \text{ W}$$

$$9) \text{ Number of neutrinos released} = \left(\frac{10^{46} \times (1.6 \times 10^{-19})^{-1}}{10^6} \right) = 6.25 \times 10^{58}$$

How many neutrinos passed through each m^2

$$N_\nu = \frac{6.25 \times 10^{58}}{4\pi (50 \times 3.09 \times 10^{19})^2} = 2 \times 10^{15}$$

taking the projected area of each human as $\sim 1 \text{ m}^2$

$$\Rightarrow 2 \times 10^{15} \text{ passed through each person}$$

$$10) \text{ Scale height, } H \sim \frac{2kT}{m_{\text{Fe}} g}$$

$$g = \frac{GM}{R^2} = \frac{6.67 \times 10^{-11} \times 1.4 \times 2 \times 10^{30}}{10^8} = 1.9 \times 10^{12} \text{ kg ms}^{-2}$$

$$m_{\text{Fe}} = 56 \times 1.67 \times 10^{-27} = 9.35 \times 10^{-26} \text{ kg} \Rightarrow H = 1.6 \times 10^{-4} \text{ m} \\ = 0.16 \text{ mm}$$