

$$\tilde{g}(k) = \frac{1}{2\pi} \int_{-\infty}^{\infty} g(y) e^{iky} dy = \frac{1}{2\pi} \int_{-\infty}^{\infty} \frac{1}{2\pi} \int_{-\infty}^{\infty} f(x) s(y-x) dx e^{iky} dy$$

$$\text{Let } z = y - x \Rightarrow y = x + z$$

$$\text{when } y = -\infty \quad z = -\infty \\ y = \infty \quad z = \infty$$

$$\begin{aligned} \tilde{g}(k) &= \frac{1}{2\pi} \int_{z=-\infty}^{\infty} \frac{1}{2\pi} \int_{x=-\infty}^{\infty} f(x) s(z) dx e^{ik(x+z)} dz \\ &= \frac{1}{2\pi} \int_{-\infty}^{\infty} f(x) e^{ikx} dx \frac{1}{2\pi} \int_{-\infty}^{\infty} s(z) e^{ikz} dz = \tilde{f}(k) \tilde{s}(k) \text{ as required.} \end{aligned}$$

The observed galaxy spectrum is a smoothed version, or convolution, of the (average) stellar spectrum, smoothed by the LOSVD.

The CCF 'falls off' from its maximum value, which should occur at \bar{v}_{los} , because features in the observed spectra no longer 'line up'.

The smaller the spread in LOSVD \Rightarrow the smaller σ_{los}

\Rightarrow the closer the galaxy spectrum is to the modelled stellar spectrum, shifted by \bar{v}_{los}

\Rightarrow the more steeply the CCF falls off for $v \neq \bar{v}_{\text{los}}$

So the width of the CCF peak is sensitive to σ_{los}

$$3) I(R) = I_0 \exp\left[-R/R_D\right]$$

$$L_D = \int_0^{2\pi} \int_0^{\infty} I(R) R dR d\theta = 2\pi I_0 \int_0^{\infty} e^{-R/R_D} R dR$$

$$\text{Let } z = R/R_D \Rightarrow L_D = 2\pi I_0 R_D^2 \int_0^{\infty} z e^{-z} dz$$

(2)

$$\int_0^{\infty} z e^{-z} dz = \left[-z e^{-z} \right]_0^{\infty} + \int_0^{\infty} e^{-z} dz$$

$$= \left[-e^{-z} \right]_0^{\infty} = 1$$

$$\Rightarrow L_D = 2\pi I_0 R_D^2$$

Fraction within α disk scale lengths:

$$f = \frac{2\pi I_0 R_D^2 \int_0^{\alpha} z e^{-z} dz}{2\pi I_0 R_D^2}$$

$$= \left[-z e^{-z} \right]_0^{\alpha} + \left[-e^{-z} \right]_0^{\alpha}$$

$$= -\alpha e^{-\alpha} + 1 - e^{-\alpha}$$

$$= 1 - (\alpha + 1) e^{-\alpha}$$

For $\alpha = 5$, $f = 1 - 6e^{-5} = 0.96$ as required

4) $M = k - 2.5 \log_{10} L$

$$= k - 2.5 \log_{10} V_{\max}^{\beta}$$

$$= k - 2.5 \beta \log_{10} V_{\max} \quad \text{as required}$$