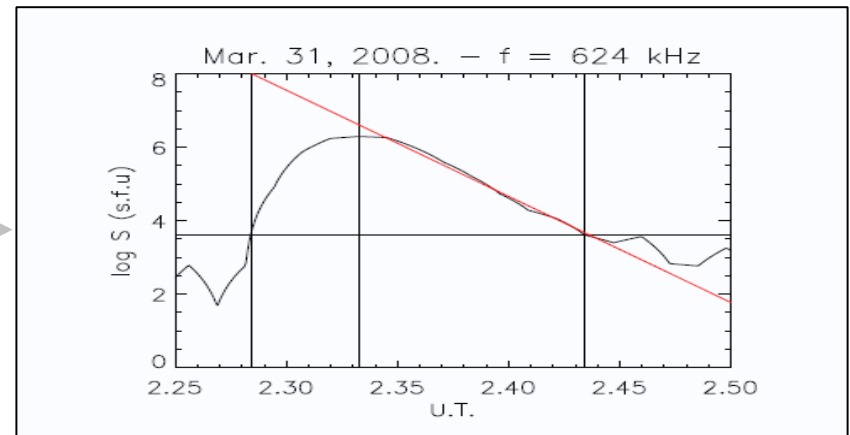
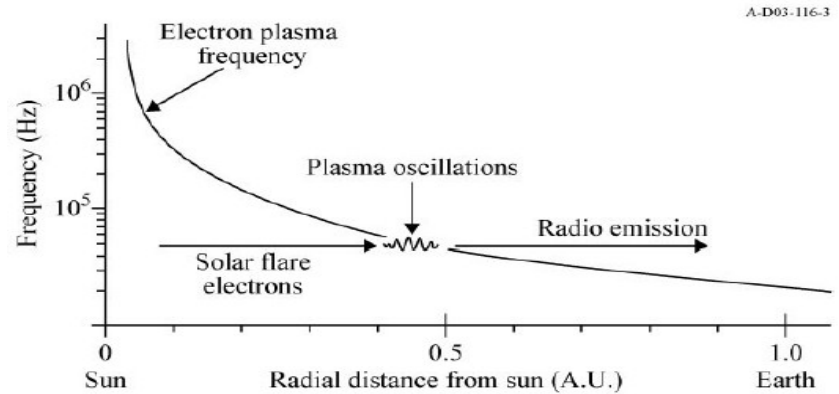
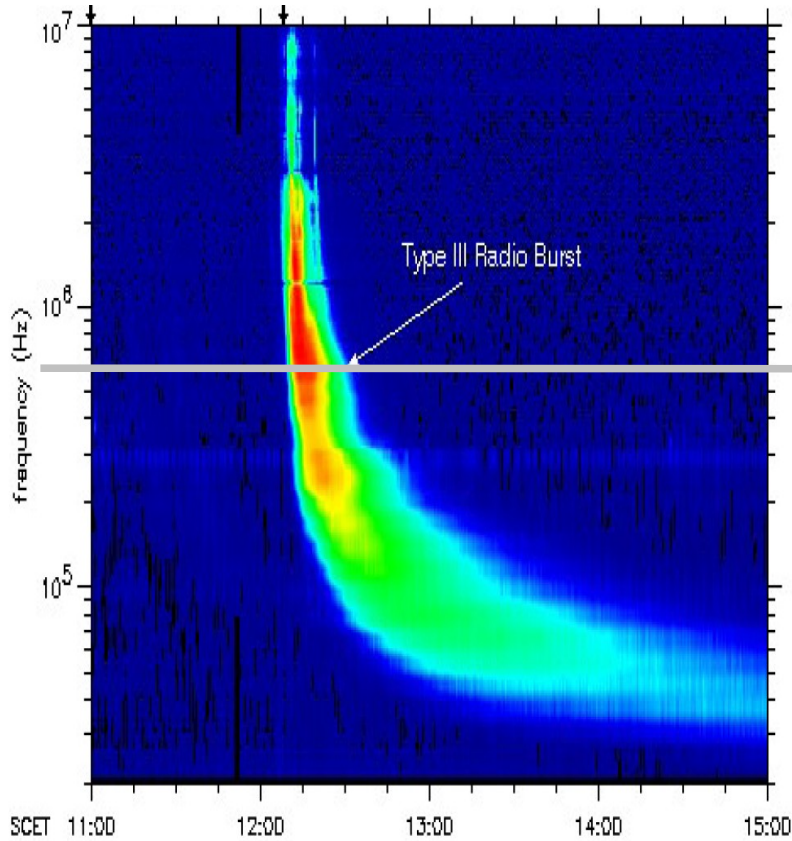


# Langmuir wave generation in the interplanetary medium

**Arnaud ZASLAVSKY**

*LESIA, Observatoire de Paris*

# Probing the interplanetary medium using radio



>2000 radio events database available in Meudon

# General questions

## 1) Langmuir wave generation

Wave-particle processes ?

## 2) Electromagnetic waves generation

Wave-wave processes ?

**WAVE-FORMS OBSERVATIONS  
+ SIMULATIONS**

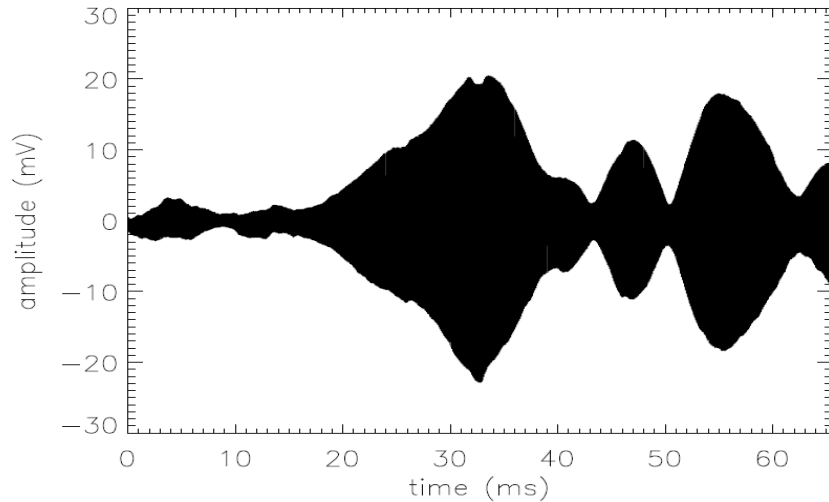
## 3) Radio waves propagation

Radiation diagrams ?

**RADIO OBSERVATIONS  
+ SIMULATIONS**

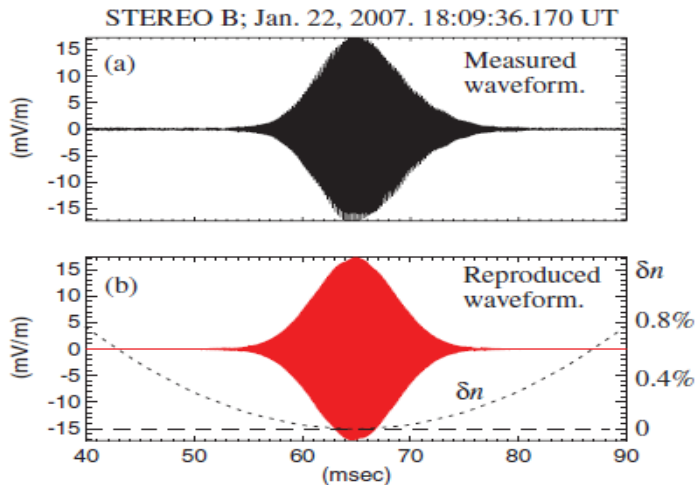
**Remote sensing of the coronal / solar wind / energetic particles**

# Electric waveforms on S/WAVES TDS



**Coherent wave-packets localized on 10-60 ms, i.e 2-15 km i.e 200-1500 electron Debye length at 1 AU**

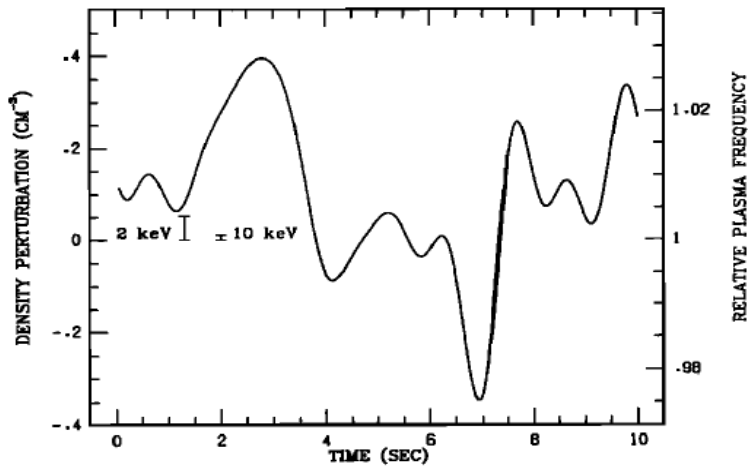
**How are such wave forms generated ?**



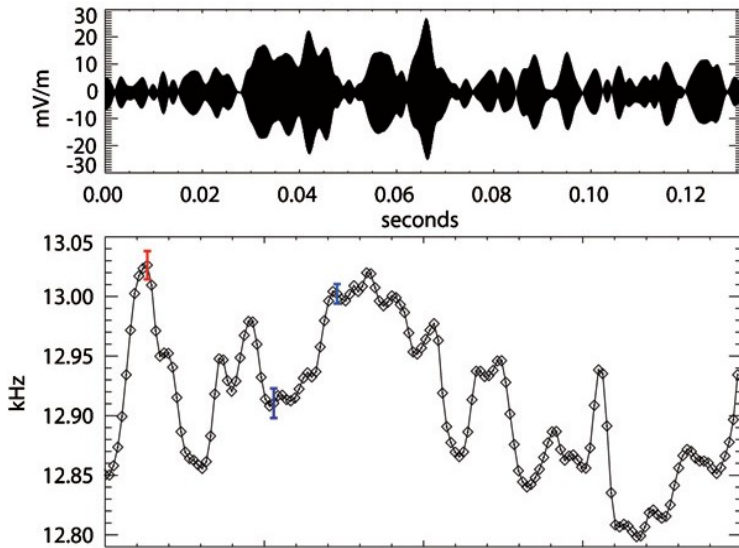
**Have to model the beam-plasma interaction in the presence of small scale density inhomogeneities affecting the waves propagation**

**Auxiliary question : Trapped eigenmode excitation ?**

# Density inhomogeneities in the solar wind

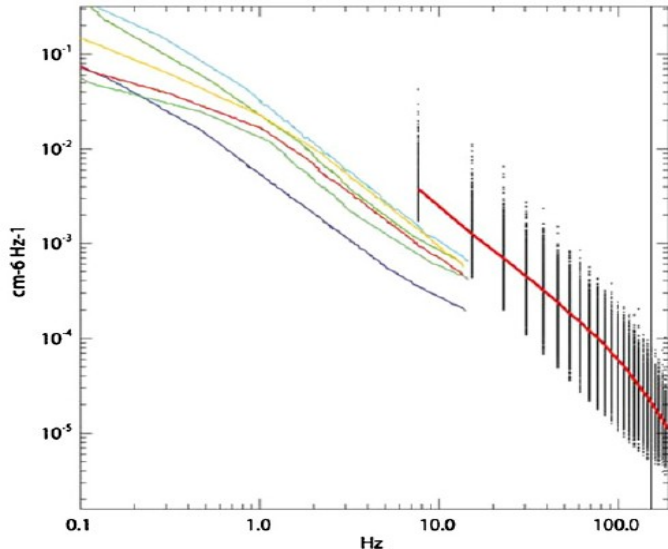


Kellogg et al, 1999



Malaspina et al, 2010

Celnikier et al, 1987 + Malaspina et al, 2010



$$\frac{dn}{n} \sim 1\% \quad k^2 l_d^2 < 1\%$$

**Turbulent density cascade, with a large enough amplitude to produce wave reflexions**

# Beam-Plasma interaction with density inhomogeneities

Propagation equation for Langmuir waves (HF Zakharov equation) :

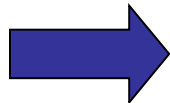
$$i \frac{\partial}{\partial t} E + \frac{3}{2} \omega_p \lambda_d^2 \nabla^2 E - \omega_p \frac{\delta n}{n_0} E = 0$$

Same equation when a external charge term is included in Poisson equation :

$$\nabla \left( i \frac{\partial}{\partial t} E + \frac{3\omega_p}{2} \lambda_d^2 \frac{\partial^2}{\partial x^2} E - \omega_p \frac{\delta n}{2n_0} E \right) = -2\pi e n_b(x, t) \omega_p e^{i\omega_p t}$$

The system is closed by an evolution equation for the beam particles :

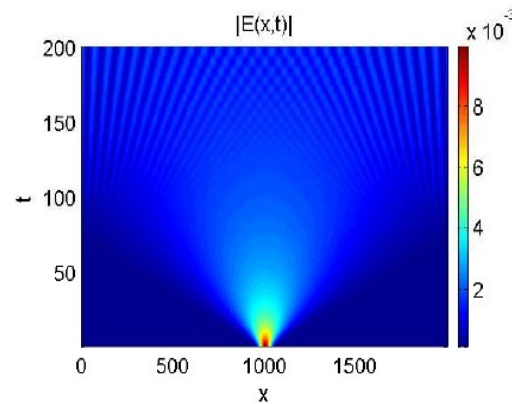
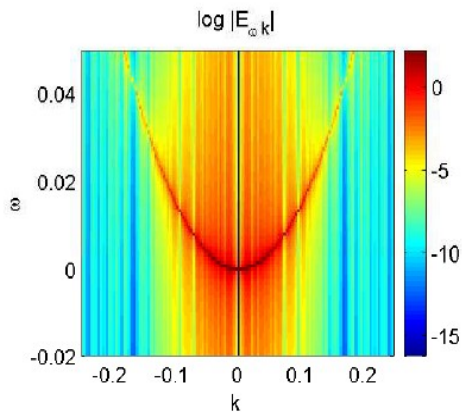
$$n_b(x, t) = \frac{\bar{n}_b L}{N} \sum_{\alpha} \delta(x - x_{\alpha}(t)), \quad d_t x_{\alpha} = v_{\alpha}, \quad d_t v_{\alpha} = -\frac{e}{m} \Re \sum E_k e^{i(kx_{\alpha} - \omega_p t)}$$



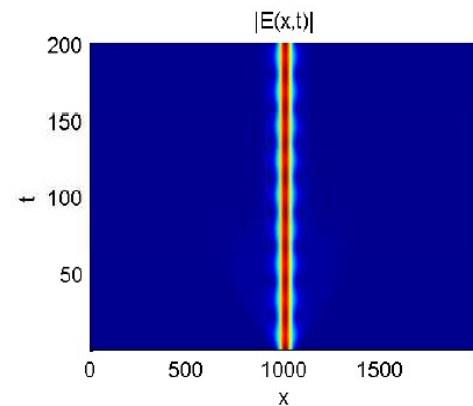
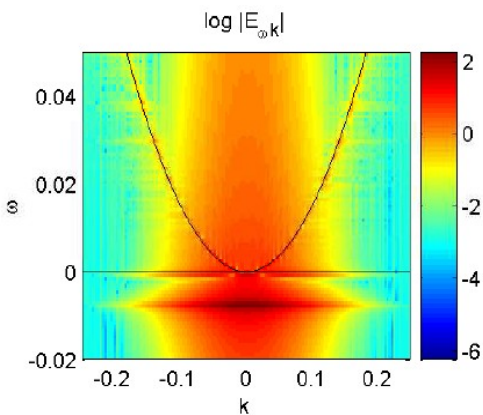
**Hybrid approach : waves treated like a fluid and beam electrons as discrete particles**

# Illustration of propagation effects (no particle beam) :

**Initial condition :** Gaussian Langmuir wave-packet launched at the center of the box

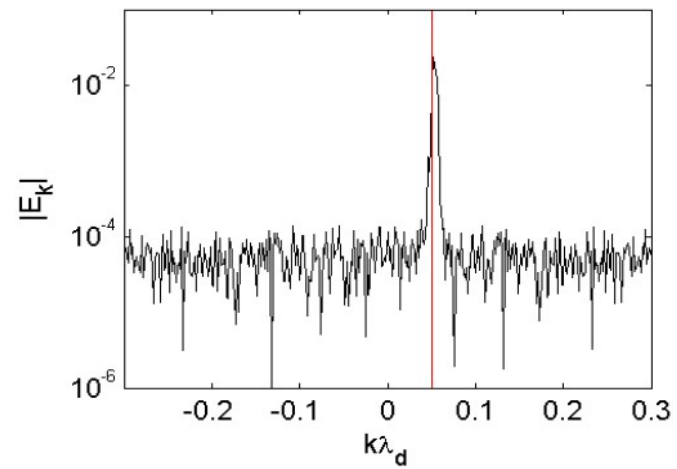
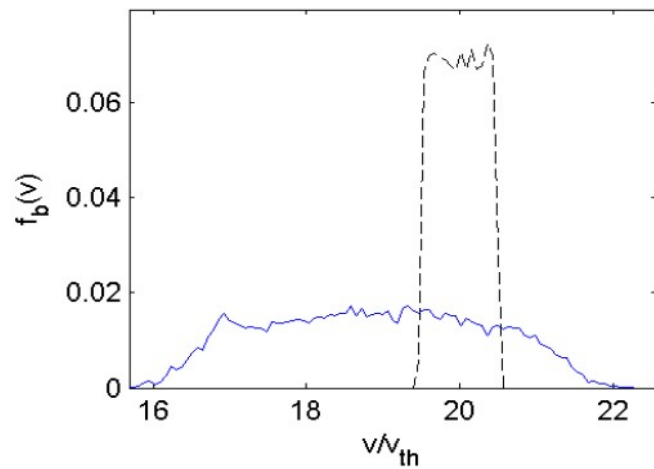
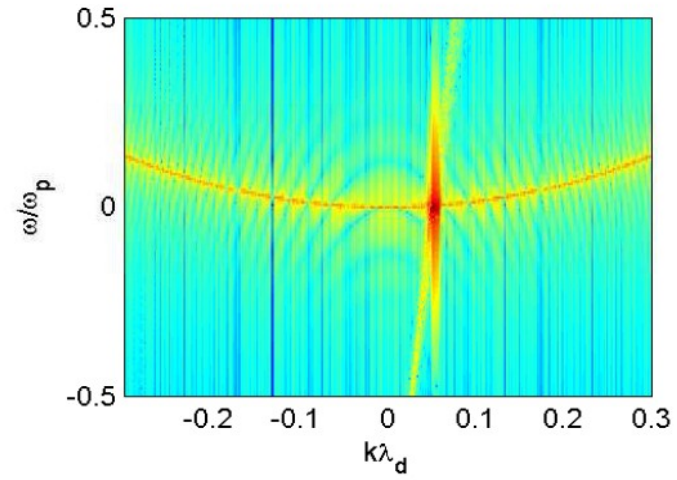
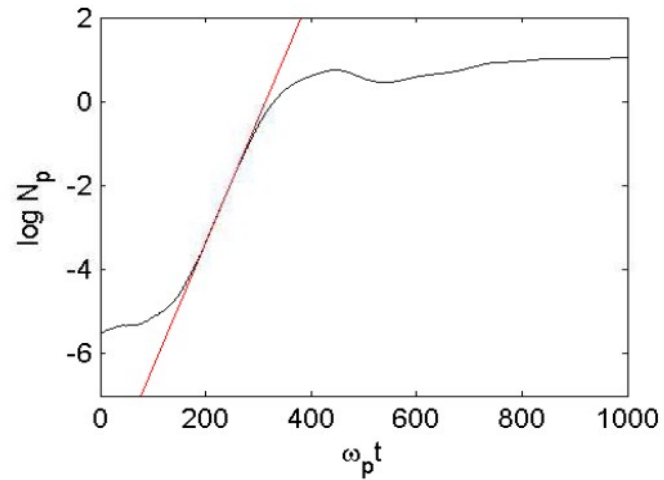


**Case 1 :** no density cavity. The waves propagate freely with the **Langmuir dispersion relation**, and we observe the **spatial dispersion of the wave packet**



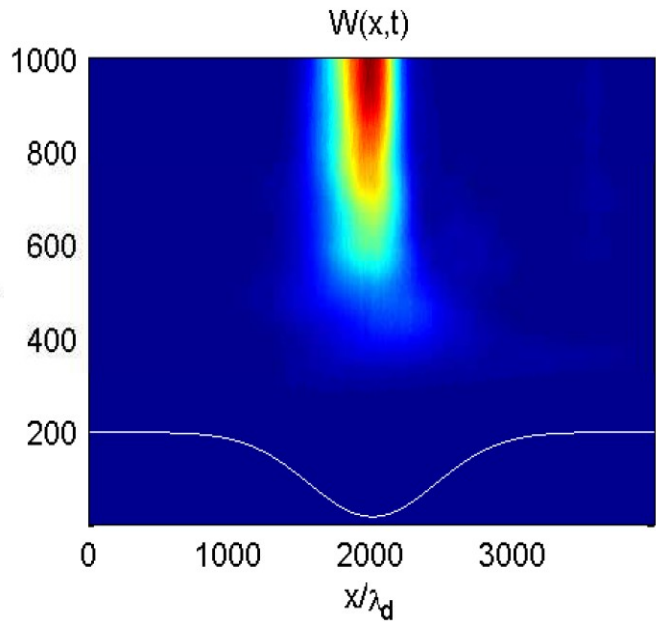
**Case 2 :** density cavity in the middle of the box. The waves propagate with a dispersion relation essentially different from the Langmuir one (negative eigen-frequency), and we observe **no spatial dispersion of the wave packet**. The wave packet is trapped.

# Beam-Plasma Instability

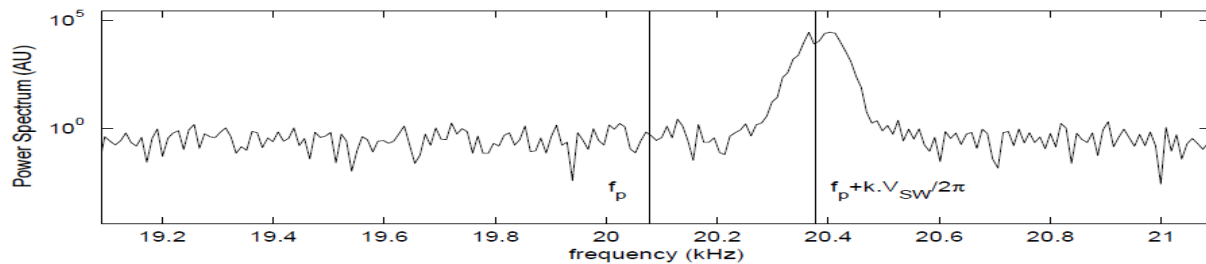
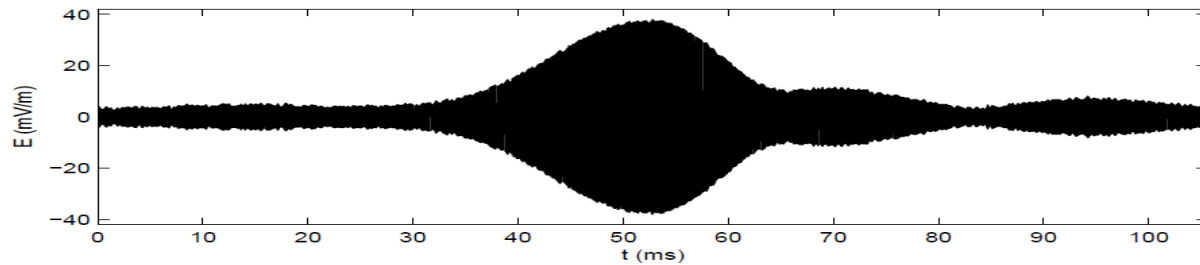




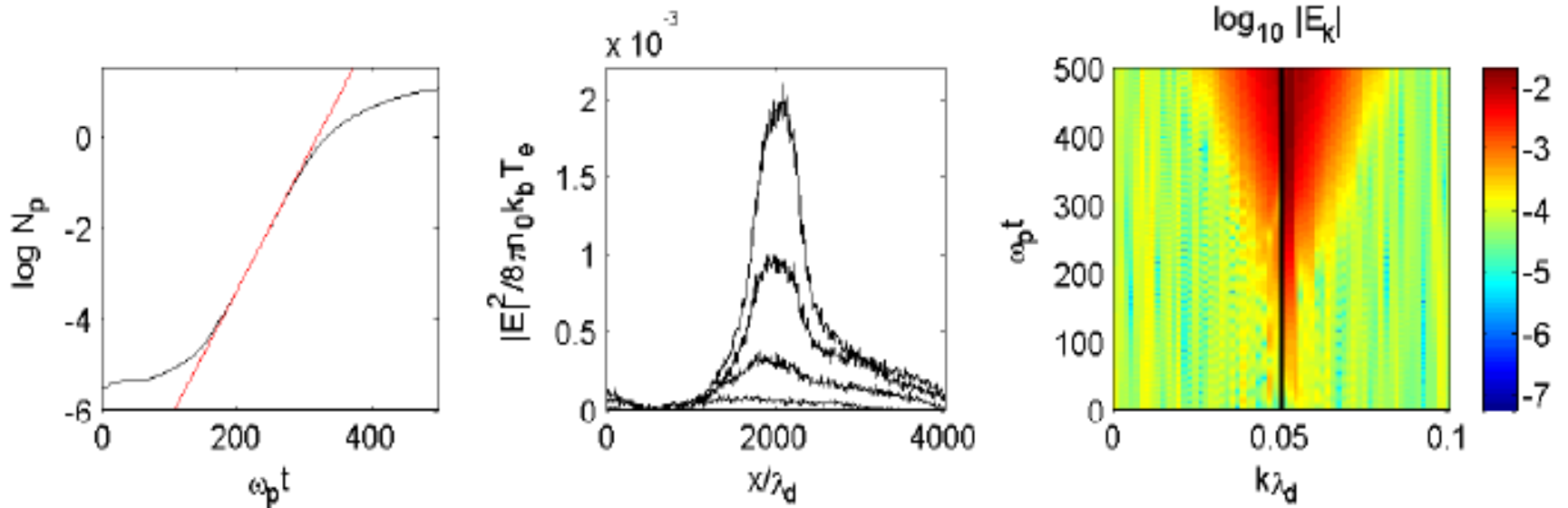
# Beam-plasma instability in the presence of a density cavity



- Beam speed :  $c/10$
- Cavity size : 1000 Debye length (10 km)
- Cavity depth : 1%



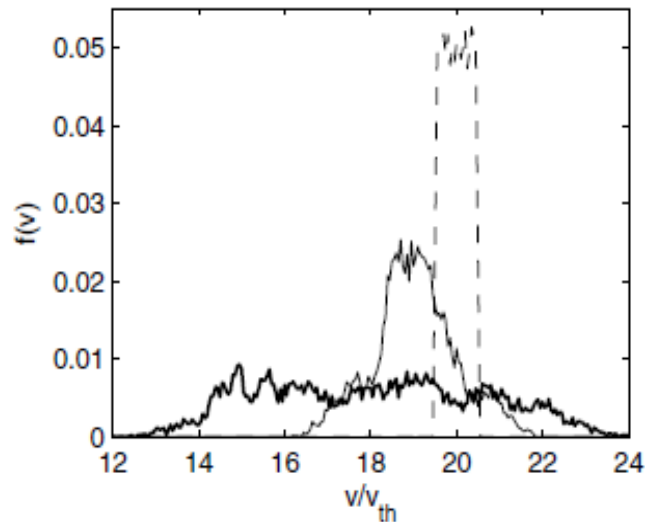
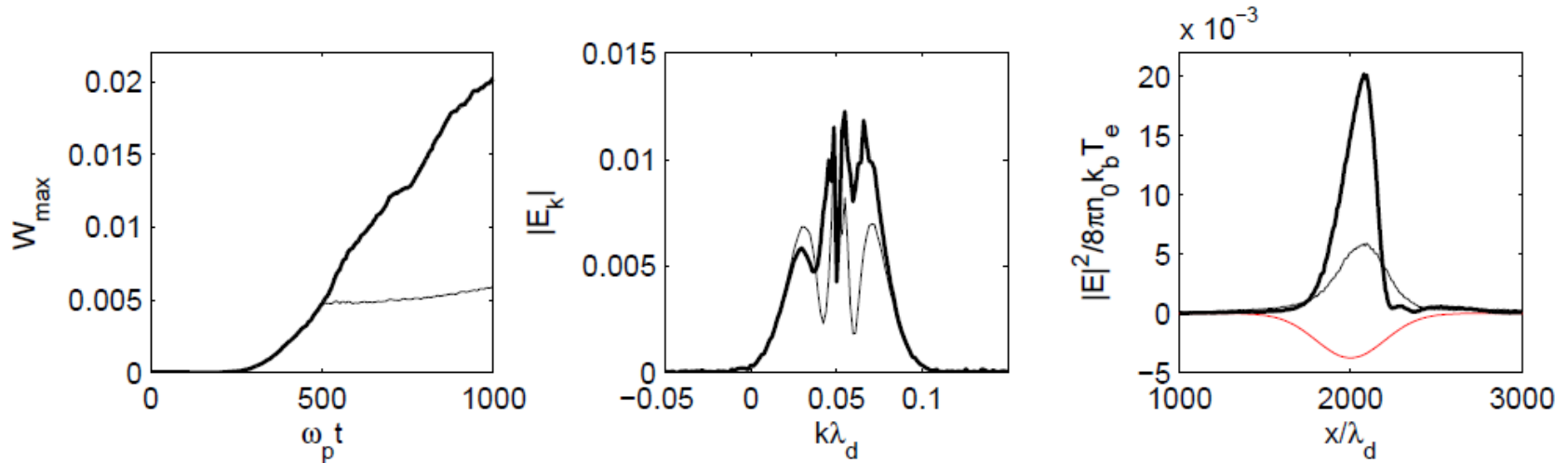
# Physics of the localization : Plasmon diffusion / reflexion



$$(\delta n E)_k = \sum_{k'+k''=k} \delta n_{k'} E_{k''} \quad \longrightarrow \quad \left( \frac{d}{dt} + i\Omega_k \right) E_k \sim -i \frac{\omega_p}{2n_0} E_{k_{res}}(t=0) \delta n_{k-k_{res}}$$

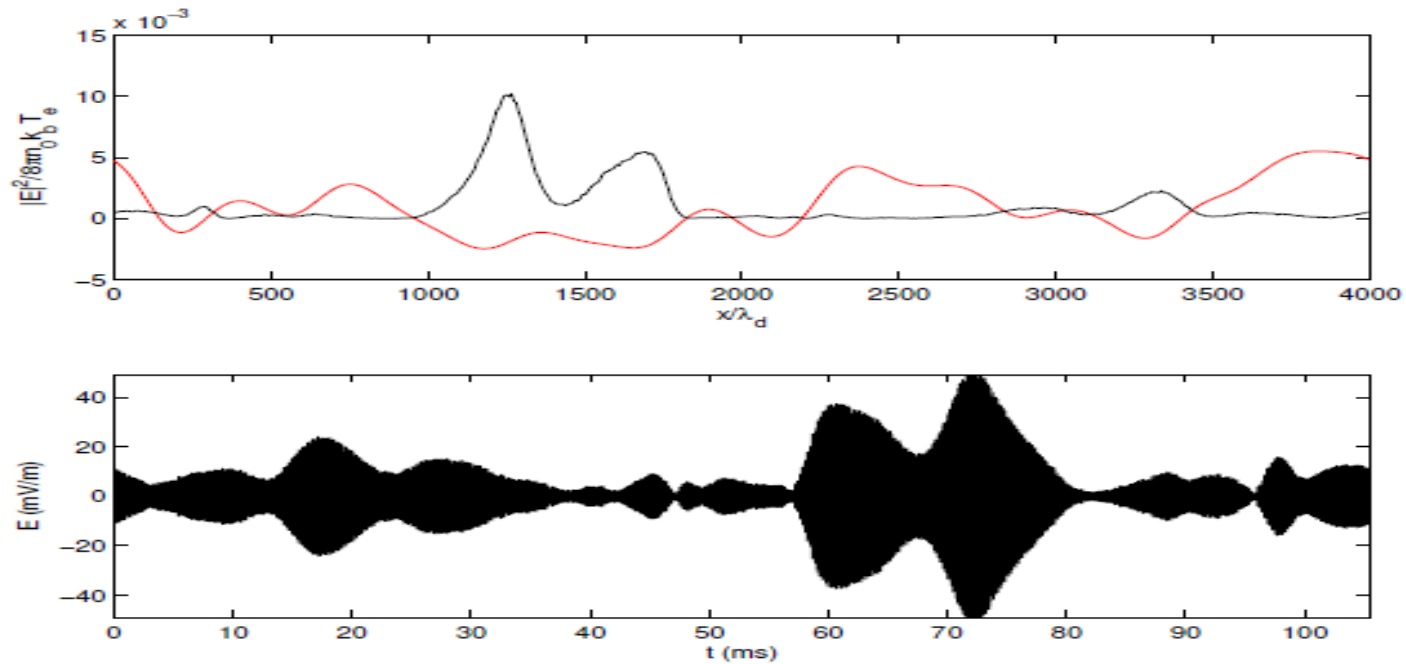
**In the linear stage of the instability, the propagation effects tends to localize the electrostatic energy in the cavity**

# Physics of the localization : wave-particle non-linear effects



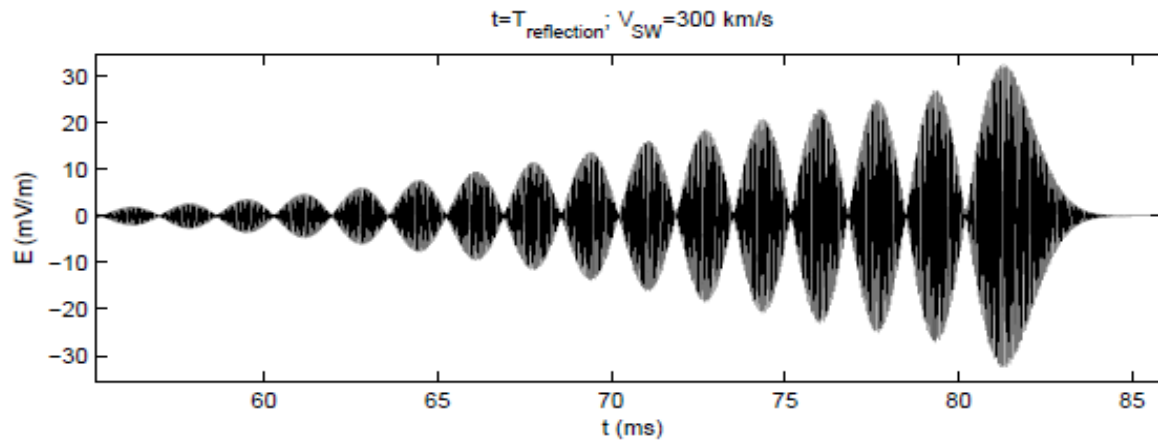
**Enhancement of the localized energy due to particles diffusion by the localized wave packet.**

## Realistic case : turbulent density spectrum

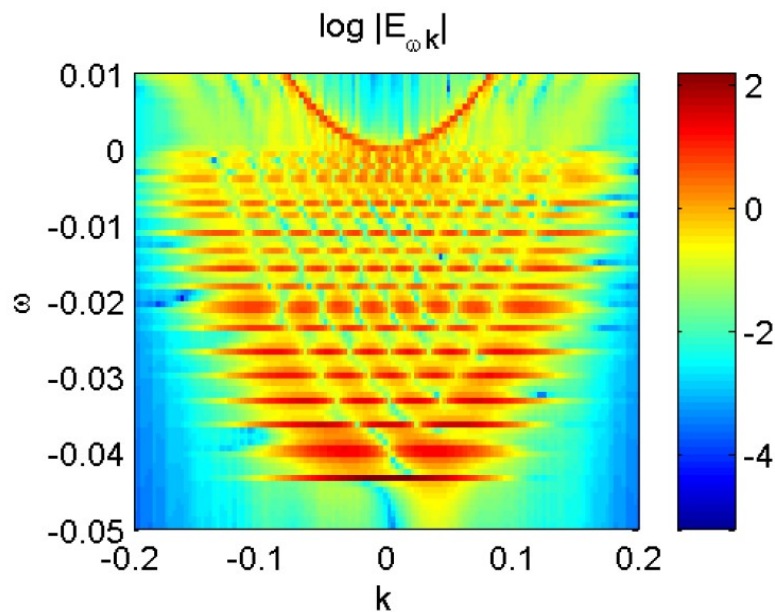


**Reproduction of wave forms similar to the one observed by S/WAVES,  
for solar wind and beam parameters similar to the one observed at 1 AU**

## Other things to study with the numerical code



**Langmuir waves near  
reflection points**



**Trapped eigen-modes :**

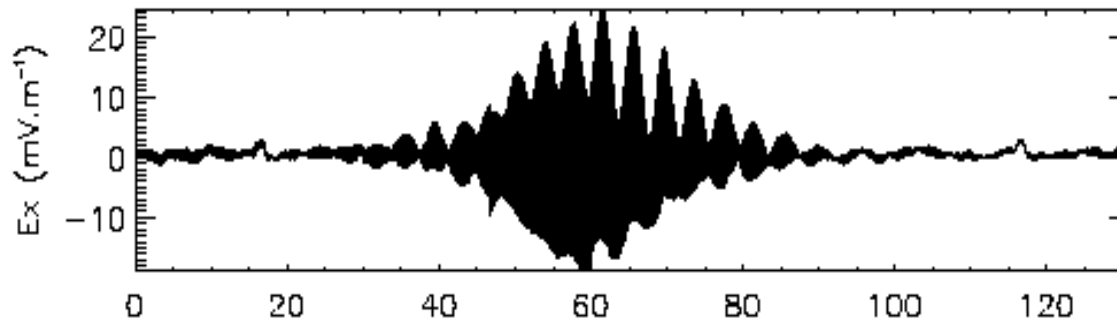
**where do they come  
from ?**

## Other things to study : radio wave generation

$$\frac{\partial^2}{\partial t^2} \delta n - c_s^2 \nabla^2 \delta n = \frac{\nabla^2 |E|^2}{16\pi m_i}$$

**Wave-wave processes :  $L \Rightarrow L' + S$**

**Observed during types III (P. Henri et al., 2009) :**



**Fondamental generation :  $L + S \Rightarrow F$**

**Harmonic generation :  $L + L' \Rightarrow H$**